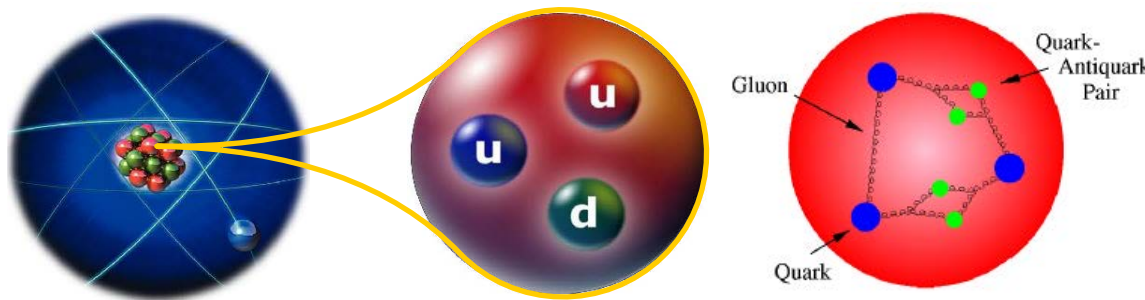


Lectures: Hans-Jürgen Wollersheim

office: 360      phone: 0188 1242294  
e-mail: h.j.wollersheim@gsi.de

Wednesday      9:55 – 10:45  
Thursday        10:50 – 11:40  
Friday            11:45 – 12:35  
Friday(T)        8:00 – 8:50

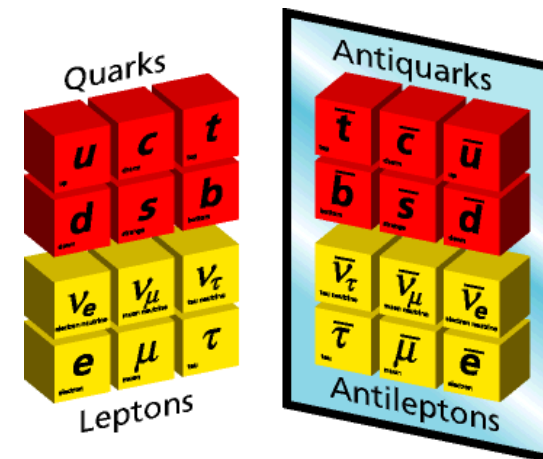
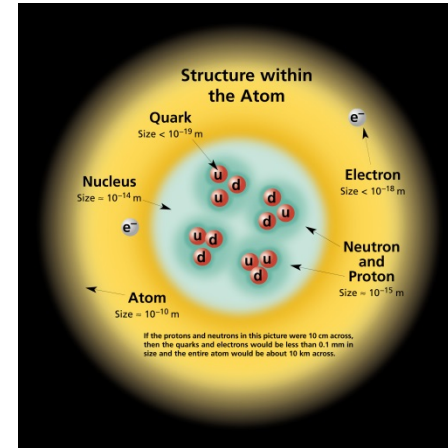
Tutor: Kanhaiya Jha ([kanhaiya.jha@iitrpr.ac.in](mailto:kanhaiya.jha@iitrpr.ac.in))

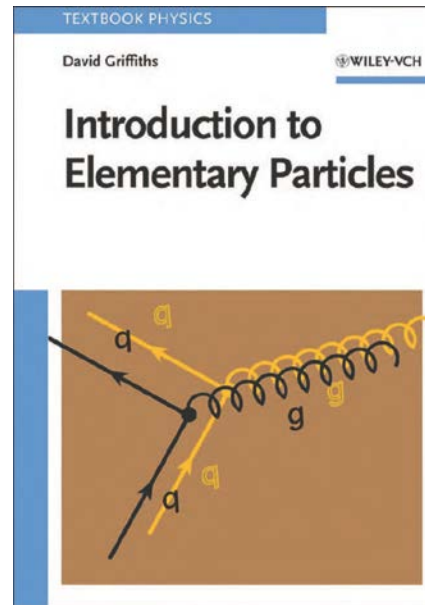
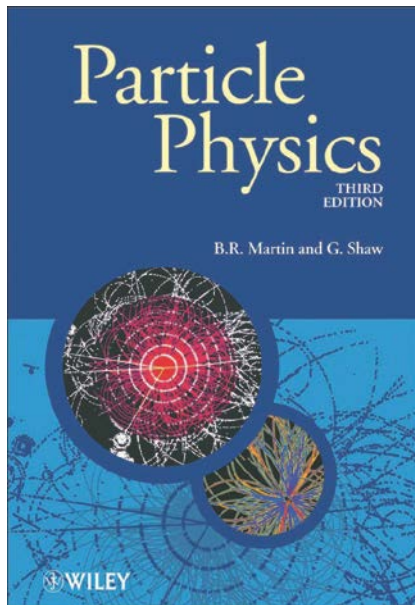


The proton does not only consist of three quarks (*blue*) being held together by gluons (*springs*), but is a sizzling place of gluons and pairs of quarks and antiquarks (*green*) interacting with each other.

# Tentative outline of Particle Physics

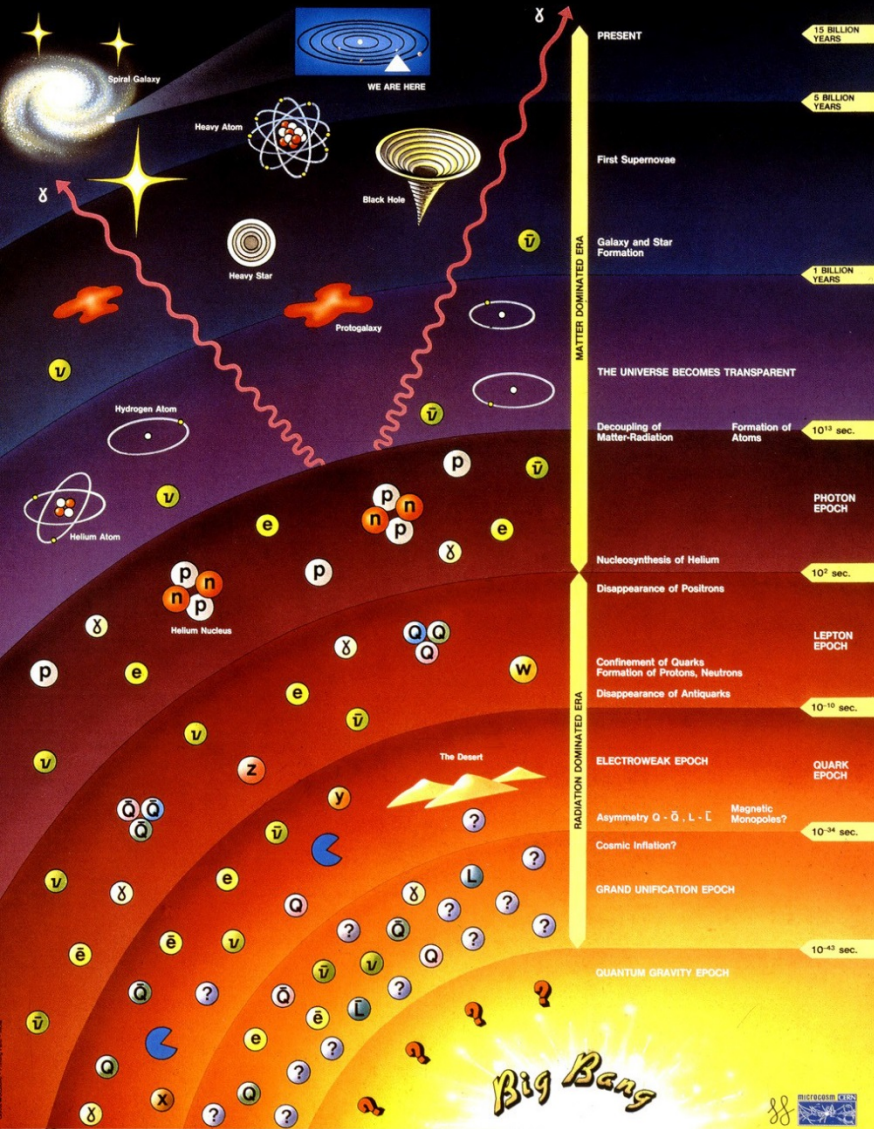
- ❖ **Introduction**
  - 100 years of particle physics
- ❖ **Basic Concepts**
  - Forces of nature
  - Standard model chart
  - Lepton and Baryon numbers
  - Units and dimensions
  - Cross section and decay rates
- ❖ **Feynman Diagrams**
  - exchange of photons
  - exchange of massive bosons
  - Yukawa potential
  - Weak, electromagnetic, strong interaction
- ❖ **Leptons**
  - Weak interaction (beta decay)
  - Electron, muon, tau
  - Tau decay
  - Lepton conservation
- ❖ **Quarks**
  - Example of baryons and meson
  - Quark number conservation
  - Isospin
  - Color a new quantum number
- ❖ **Summary & Open Questions**
- ❖ **Beyond the Standard Model**





❖ Recommended Textbook

## History of the Universe



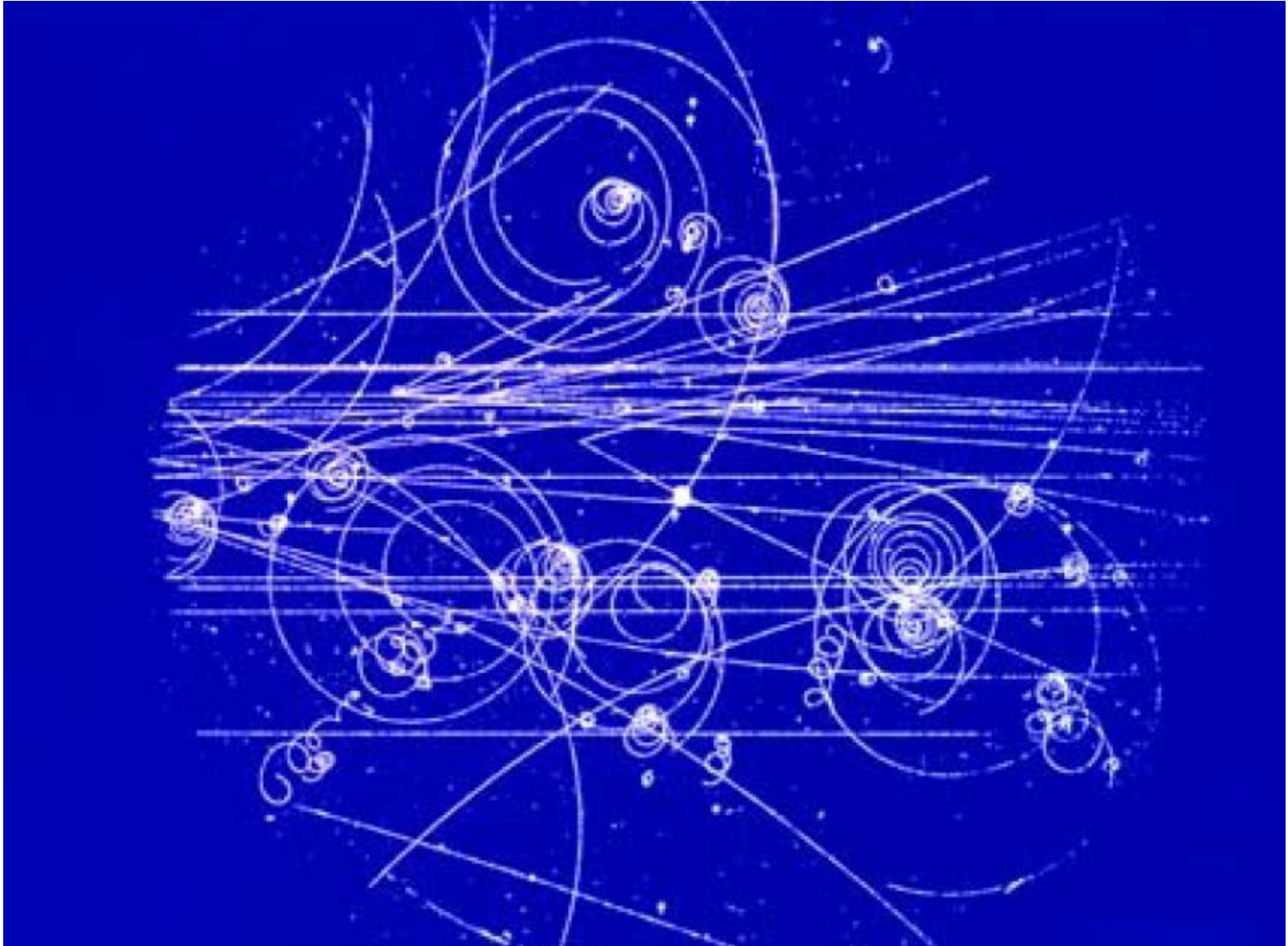
era of gravitation

10<sup>13</sup> s

era of particle physics



# 100 years of particle physics



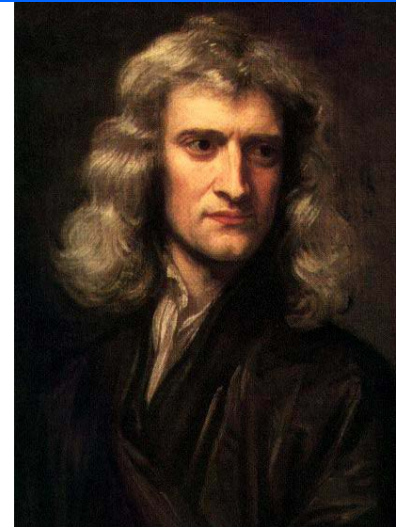
# Corpuscular Theory of Light

Light consist out of particles (Newton)



Light is a wave (Huygens)

- ❖ Mainly because of Newton's prestige, the corpuscle theory was widely accepted (more than 100 years)
- ❖ Failing to describe interference, diffraction, and polarization (e.g. Fresnel) corpuscle theory was abandoned for Huygens wave theory
- ❖ Wave theory strongly supported by Maxwell equations and by H. Hertz experiments
- ❖ Until in the early 20<sup>th</sup> century ...



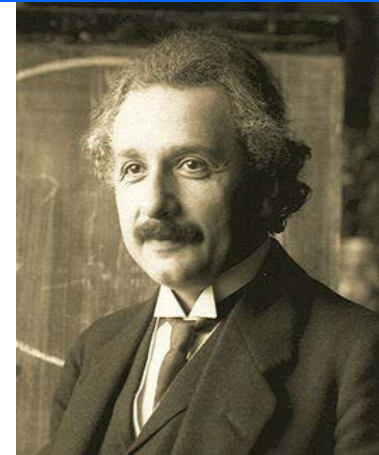
Isaac Newton  
1643-1727



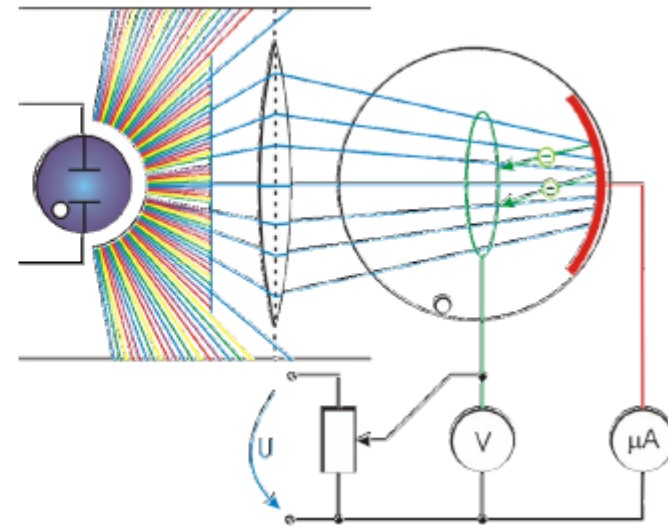
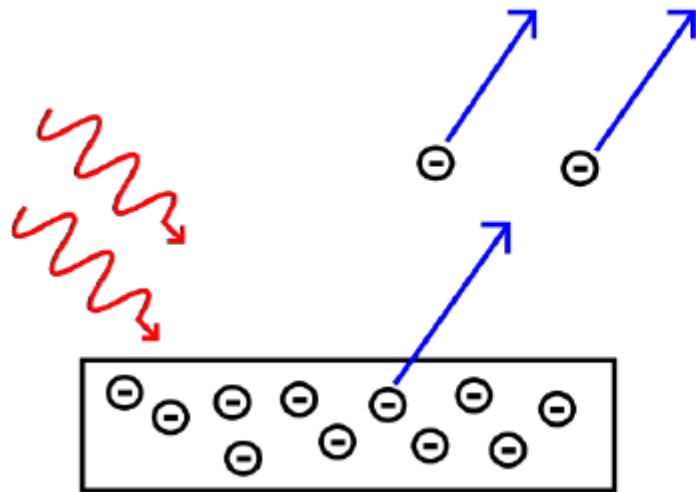
Christiaan Huygens  
1629-1695

# Photoelectric Effect

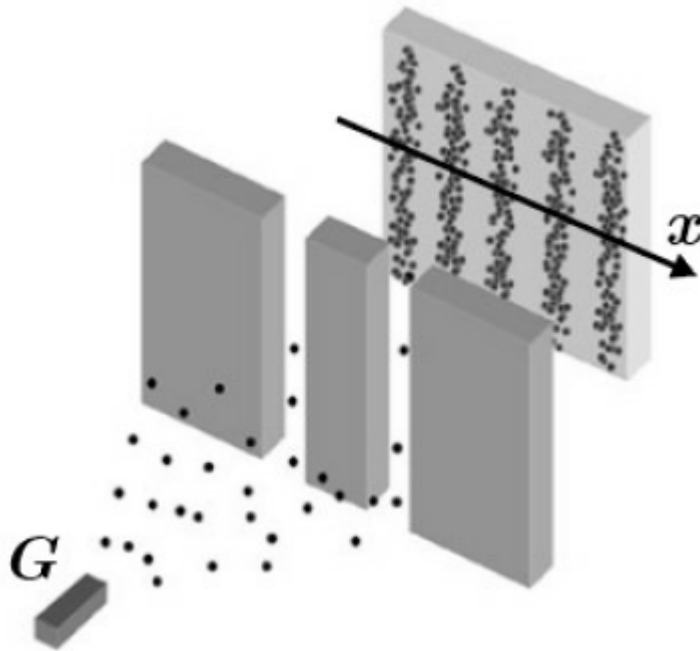
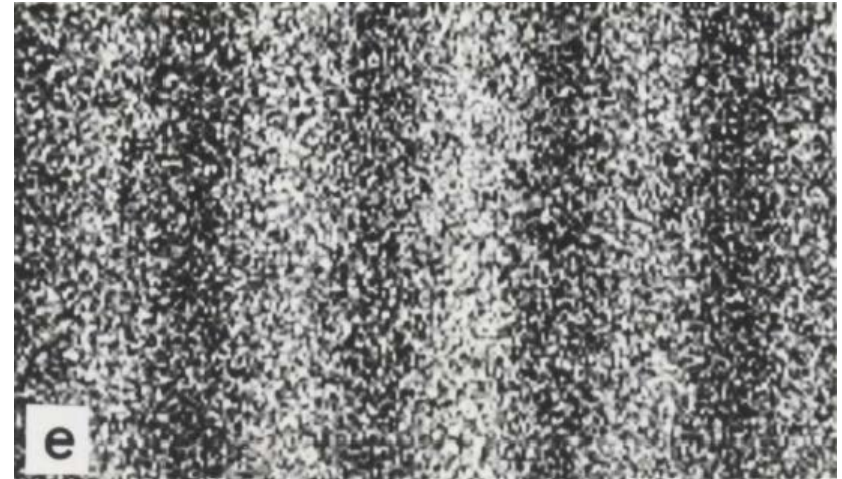
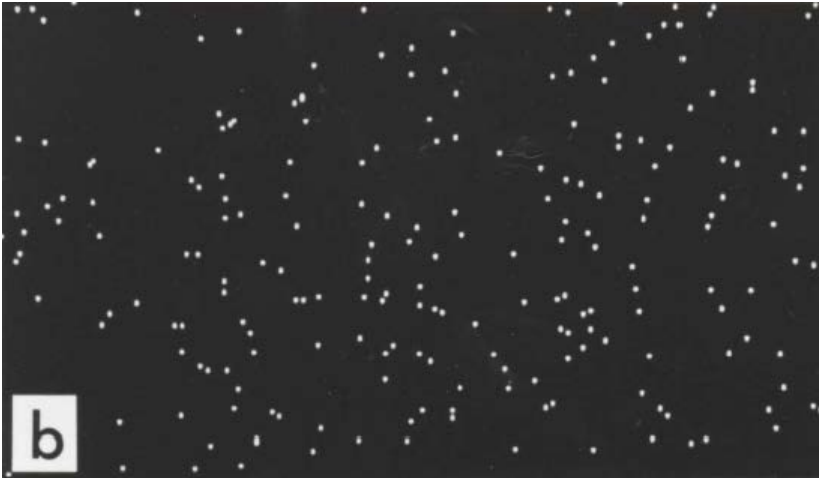
- ❖ **Observation:** 1836 Becquerel  
Metal absorbs light and emits electrons →  
Maximum energy of electrons independent on intensity  
(number of photons)
- ❖ **Interpretation:** 1905 Einstein (Nobel prize 1921)  
Light consists of particles (photons) with quantized energy



Albert Einstein  
1879-1955



# Double Slit with Electrons



Electron (particles) have wavelength

*De Broglie wavelength:*

$$\lambda = \frac{h}{p} = \frac{\text{Planck constant}}{\text{momentum (= mass} \cdot \text{velocity)}}$$



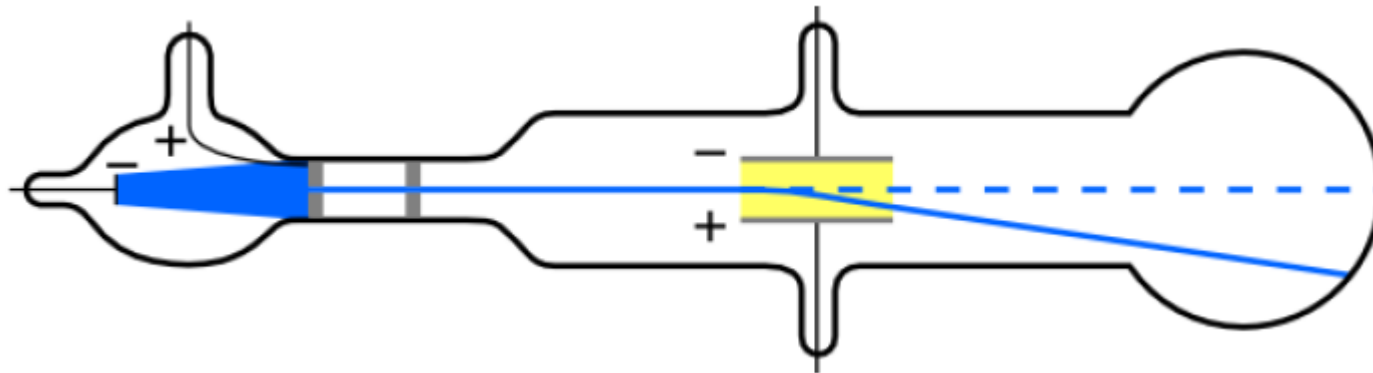
# The Electron

## ❖ **Observation:** 1897 (Nobel Prize 1906)

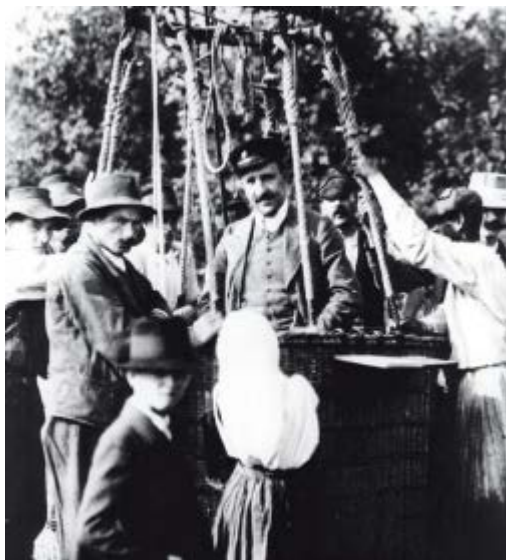
- Constituents of cathode rays deflected by electric field
- Constituents of cathode rays deflected by magnetic field + heating of thermal junction → first mass/charge ratio
- Higher precision of mass/charge from comparing deflection by electric and magnetic fields



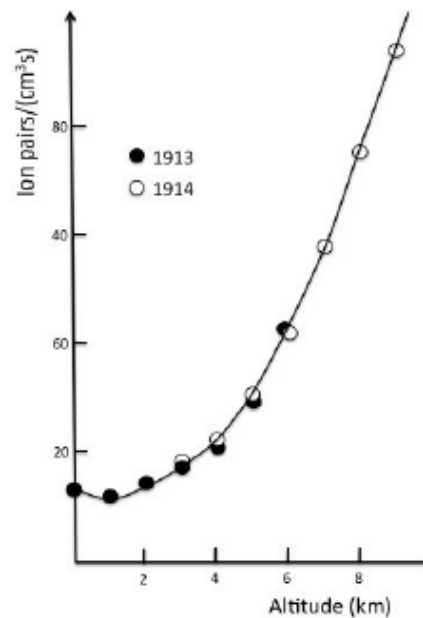
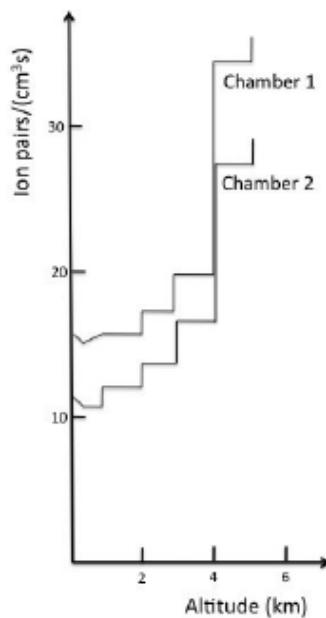
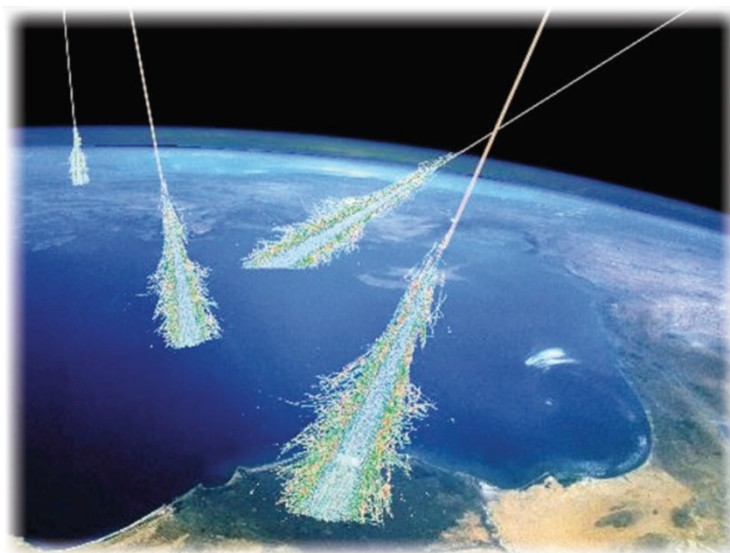
Joseph J. Thomson  
1856-1940



# Ionization Measurements

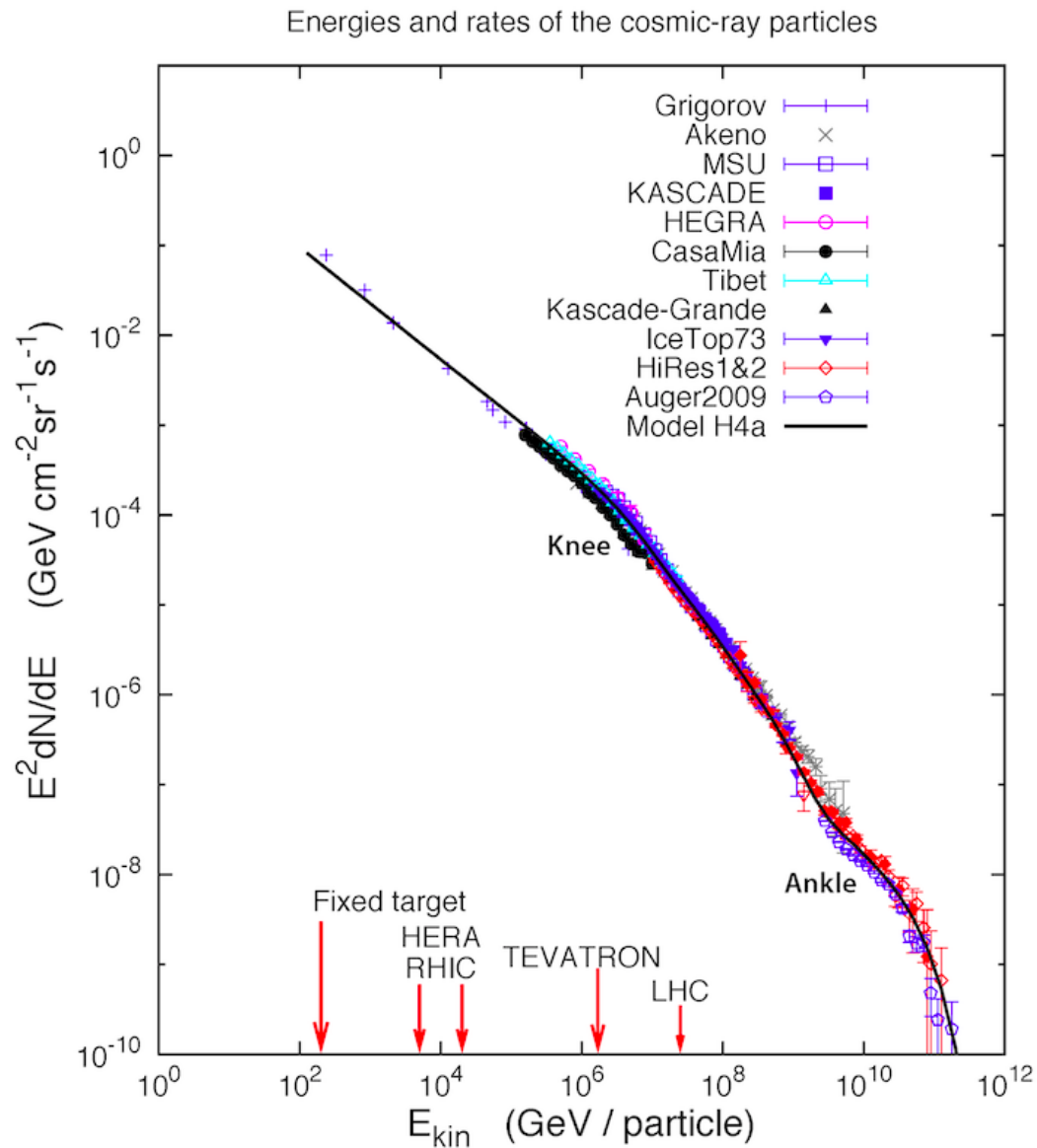


Victor Hess (1912)  
discovery of cosmic showers



Nobel Prize 1936

# The Cosmic Accelerator

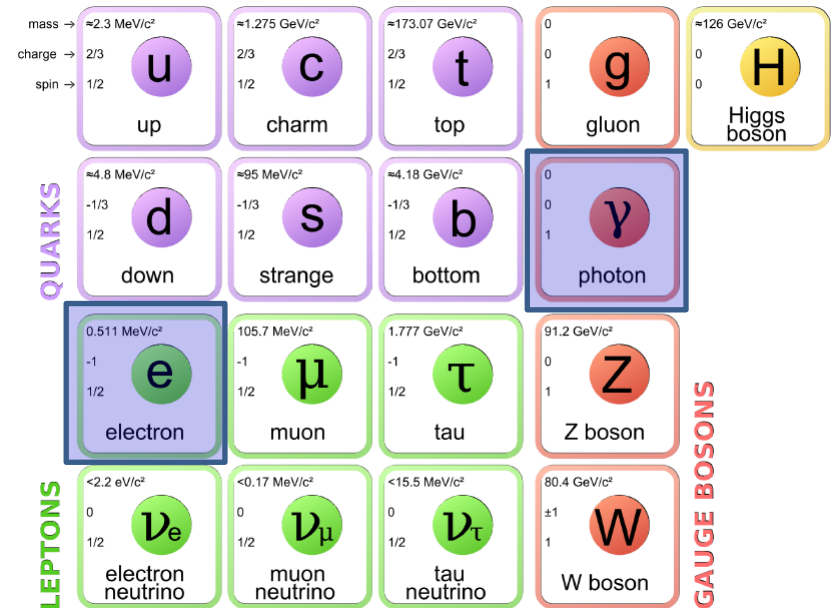


1912

~100 years  
→

2017

- ❖ electron
- ❖ photons
- ❖ nuclei (protons)





- ❖ electron
- ❖ photons
- ❖ nuclei (protons)

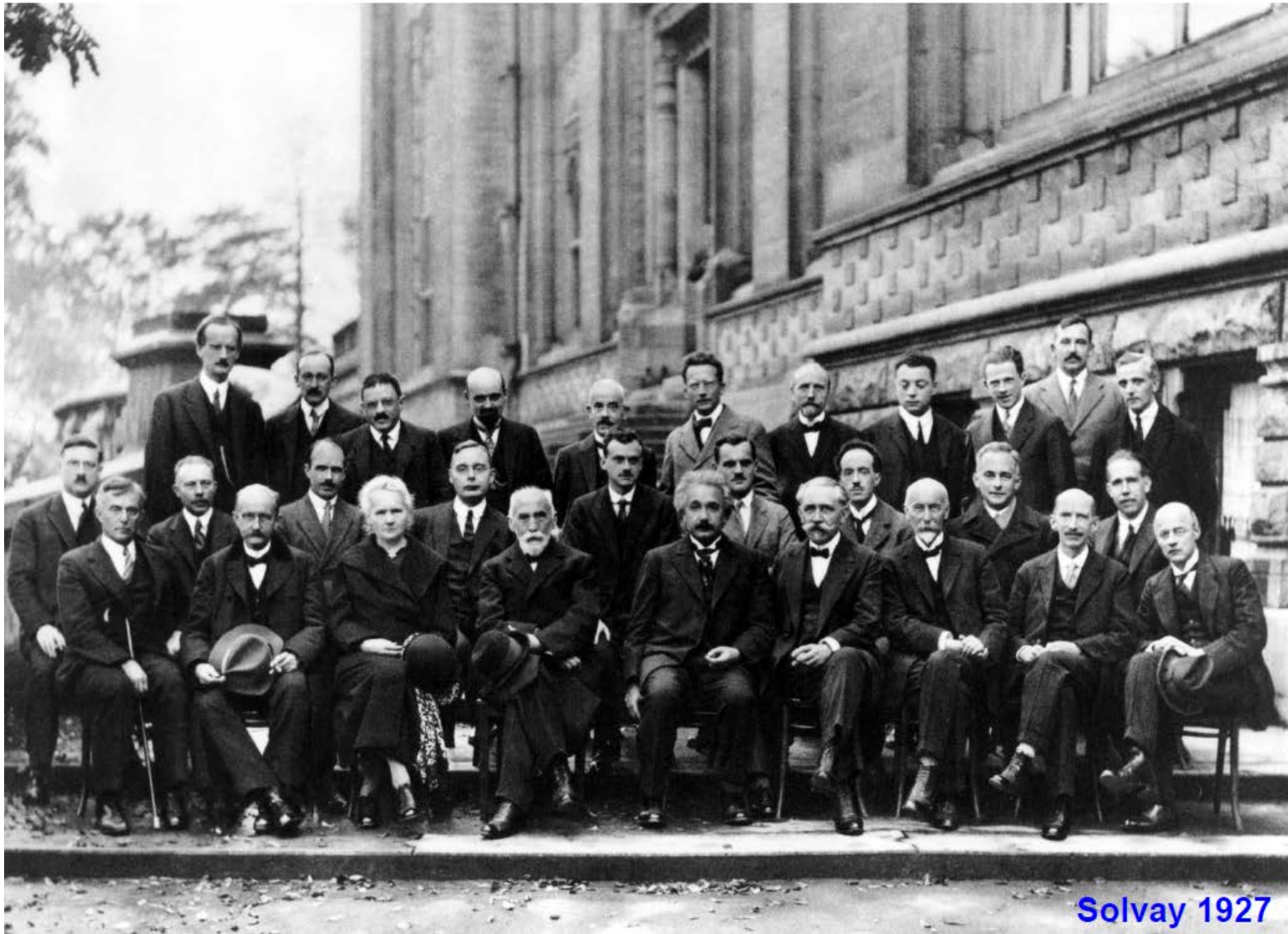
Standard Model  
Lagrangian

- 1 is ultra specific to the gluon, the boson that carries the strong force
  - 2 is dedicated to explain interactions between bosons, particularly W and Z
  - 3 describes how elementary matter particles interact with the weak force
  - 4 describes how matter particles interact with Higgs ghosts, virtual artifacts
- In quantum mechanics, there is no single path a particle can take

$$\begin{aligned}
 & -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \\
 & \frac{1}{2}ig_s^2(\bar{q}_i^\sigma \gamma^\mu q_j^\sigma)g_\mu^a + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
 & M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
 & \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2} M \phi^0 \phi^0 - \beta_h [\frac{2M^2}{g^2} + \\
 & \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-)] + \frac{2M^4}{g^2} \alpha_h - ig_{c_w} [\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\mu W_\nu^- - \\
 & W_\nu^- \partial_\mu W_\mu^+) - ig_{s_w} [\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\mu W_\mu^- - \\
 & W_\mu^- \partial_\mu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
 & \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^+ W_\nu^- + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\nu^+ W_\nu^-) + \\
 & g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
 & \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
 & g M W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig [W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
 & W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g [W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H \partial_\mu \phi^+ - \\
 & \phi^+ \partial_\mu H)] + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig \frac{s_w^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
 & ig_{s_w} M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
 & ig_{s_w} A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
 & \frac{1}{4}g^2 \frac{1}{c_w} Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\
 & g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda - \\
 & \bar{d}_j^\lambda (\gamma \partial + m_d^\lambda) d_j^\lambda + ig_{s_w} A_\mu [-(\bar{e}^\lambda \gamma^\mu e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)] + \\
 & \frac{ig}{4c_w} Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - \\
 & 1 - \gamma^5) u_j^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}} W_\mu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + \\
 & (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}} W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \\
 & \gamma^5) u_j^\lambda)] + \frac{ig}{2\sqrt{2}} \frac{m_\lambda^\lambda}{M} [-\phi^+ (\bar{\nu}^\lambda (1 - \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \\
 & \frac{g}{2} \frac{m_\lambda^\lambda}{M} [H (\bar{e}^\lambda e^\lambda) + i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] + \frac{ig}{2M\sqrt{2}} \phi^+ [-m_d^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + \\
 & m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa)] + \frac{ig}{2M\sqrt{2}} \phi^- [m_d^\lambda (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\lambda (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger (1 - \\
 & \gamma^5) u_j^\kappa) - \frac{g}{2} \frac{m_\lambda^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2} \frac{m_\lambda^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_\lambda^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \\
 & \frac{ig}{2} \frac{m_\lambda^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \\
 & \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + ig_{c_w} W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + ig_{s_w} W_\mu^+ (\partial_\mu \bar{Y} X^- - \\
 & \partial_\mu \bar{X}^+ Y) + ig_{c_w} W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + ig_{s_w} W_\mu^- (\partial_\mu \bar{X}^- Y - \\
 & \partial_\mu \bar{Y} X^+) + ig_{c_w} Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + ig_{s_w} A_\mu (\partial_\mu \bar{X}^+ X^+ - \\
 & \partial_\mu \bar{X}^- X^-) - \frac{1}{2}g M [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w} \bar{X}^0 X^0 H] + \\
 & \frac{1-2c_w^2}{2c_w} ig M [\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w} ig M [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \\
 & ig M s_w [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \frac{1}{2}ig M [\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
 \end{aligned}$$



# Quantum Mechanics



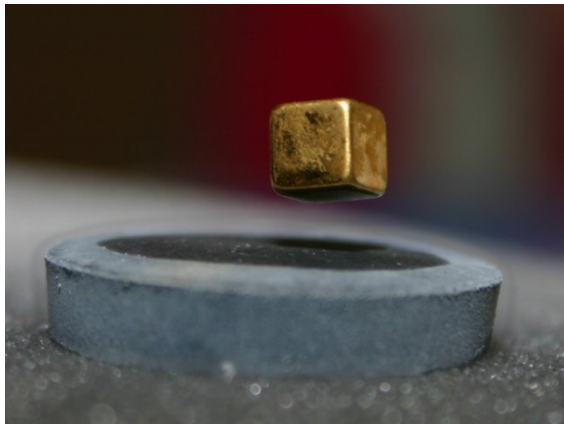
A. Piccard, E. Henriot, P. Ehrenfest, E. Herzen, Th. de Donder, E. Schrödinger, J.E. Verschaffelt, W. Pauli, W. Heisenberg, R.H. Fowler, L. Brillouin  
P. Debye, M. Knudsen, W.L. Bragg, H.A. Kramers, P.A.M. Dirac, A.H. Compton, L. de Broglie, M. Born, N. Bohr,  
I. Langmuir, M. Planck, M. Curie, H.A. Lorentz, A. Einstein, P. Langevin, Ch.-E. Guye, C.T.R. Wilson, O.W. Richardson

# Fermions and Bosons

Fundamental characteristics of particles: *Spin* (“self angular momentum”)

- ❖ Integer values ( $0, \pm 1, \dots$ ) → *Bosons*
- ❖ Half integer values ( $\pm 1/2, \pm 3/2, \dots$ ) → *Fermions*

*Bosons* (Cooper pairs ...) can be described by *common wave function*  
→ funny effects (super conductivity, super fluidity, ...)



*Fermions* (electrons or protons ...) must be in different states  
→ *Pauli's exclusion principle* (basis of all chemistry...)



Satyendranath Bose



Enrico Fermi



Wolfgang Pauli

# Matter Waves for Particles without Spin

Non relativistic:

Kinematics:

$$E = \frac{\vec{p}^2}{2m}$$

Quantum Mechanics:

$$E \rightarrow i\hbar \frac{\partial}{\partial t} \quad \text{and} \quad \vec{p} \rightarrow -i\hbar \vec{\nabla}$$

Wave Equation:

$$i\hbar \frac{\partial}{\partial t} \psi = \frac{-\hbar^2}{2m} \nabla^2 \psi$$

Relativistic:

$$E^2 = p^2 c^2 + m^2 c^4$$

$$E \rightarrow i \frac{\partial}{\partial t} \quad \text{and} \quad \vec{p} \rightarrow -i \vec{\nabla}$$

$$-\frac{1}{c^2} \frac{\partial^2}{\partial t^2} \phi = -\nabla^2 \phi + \frac{m^2 c^2}{\hbar^2} \phi$$

Klein-Gordon equation

density:

$$\rho(x, t) = |\psi(x, t)|^2$$

continuity equation:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot j = 0$$

Schrödinger equation:

$$\frac{\partial}{\partial t} \psi = \frac{i\hbar}{2m} \nabla^2 \psi$$

continuity equation:

$$\frac{\partial}{\partial t} (\psi^* \psi) = -\nabla \cdot \left[ \frac{i\hbar}{2m} (\psi \nabla \psi^* - \psi^* \nabla \psi) \right]$$

plane waves:

$$\psi = N \cdot e^{i(p \cdot x - E \cdot t)/\hbar} \quad E = p^2 / 2m$$

are solutions to the free  
Schrödinger equation

$\mathbf{p} \rightarrow -\mathbf{p}$  gives to same result



# Matter Waves for Particles without Spin

## Non relativistic:

Kinematics:

$$E = \frac{\vec{p}^2}{2m}$$

Quantum Mechanics:

$$E \rightarrow i\hbar \frac{\partial}{\partial t} \quad \text{and} \quad \vec{p} \rightarrow -i\hbar \vec{\nabla}$$

Wave Equation:

$$i\hbar \frac{\partial}{\partial t} \psi = \frac{-\hbar^2}{2m} \nabla^2 \psi$$

## Relativistic:

$$E^2 = p^2 c^2 + m^2 c^4$$

$$E \rightarrow i \frac{\partial}{\partial t} \quad \text{and} \quad \vec{p} \rightarrow -i \vec{\nabla}$$

$$-\frac{1}{c^2} \frac{\partial^2}{\partial t^2} \phi = -\nabla^2 \phi + \frac{m^2 c^2}{\hbar^2} \phi$$

Klein-Gordon equation

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2 + m^2 \right) \phi(x) = 0$$

$$\phi(x) = N e^{i(p \cdot x - E \cdot t)} = e^{-i p_\mu \cdot x^\mu}$$

Unlike the Schrödinger equation, the Klein-Gordon equation does not contain factor  $i$ . Consequently, it can have both real and complex solutions.

plane waves:

satisfies the dispersion relation

$$E^2 = p^2 + m^2$$

continuity equation:

The factor  $i$  is introduced to make the density real

$$\frac{\partial}{\partial t} i \left( \phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right) = \nabla \cdot [i(\phi^* \nabla \phi - \phi \nabla \phi^*)]$$

$$\rho = 2|N|^2 E$$

$$j = 2|N|^2 p$$

# Matter Waves for Particles with Spin

$$\left( \beta mc^2 + \sum_{k=1}^3 \alpha_k p_k c \right) \psi(x, t) = i\hbar \frac{\partial \psi(x, t)}{\partial t}$$

Dirac equation



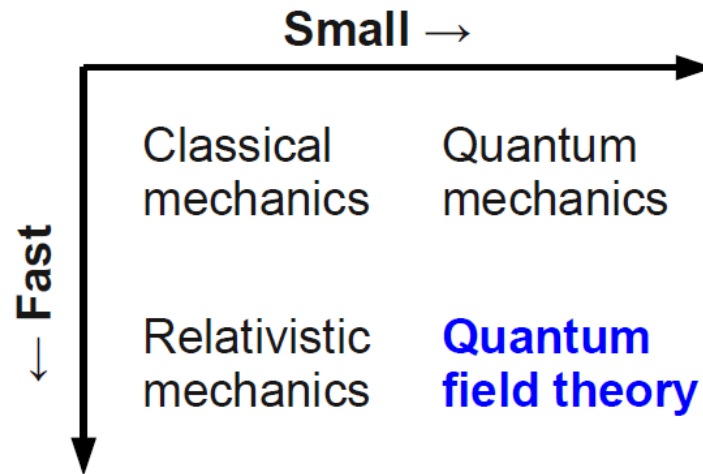
Paul Dirac 1902-1984

The problem with the Klein-Gordon equation: it is second order in derivatives. In 1928, Dirac found the first-order form having the same solutions.  $\alpha_k$  and  $\beta$  are 4x4 matrices and  $\Psi$  are four-component wave functions: *spinors* (for particles with spin  $\frac{1}{2}$ ).

$$i\hbar \gamma^\mu \partial_\mu \psi - mc\psi = 0$$

$$\gamma^0 = \begin{pmatrix} I_2 & 0 \\ 0 & I_2 \end{pmatrix} \quad \gamma^1 = \begin{pmatrix} 0 & \sigma_x \\ -\sigma_x & 0 \end{pmatrix} \quad \gamma^2 = \begin{pmatrix} 0 & \sigma_y \\ -\sigma_y & 0 \end{pmatrix} \quad \gamma^3 = \begin{pmatrix} 0 & \sigma_z \\ -\sigma_z & 0 \end{pmatrix}$$

# Quantum Field Theory

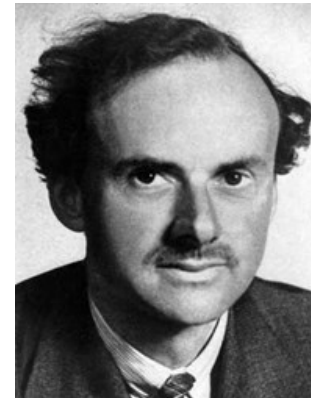


- ❖ **First major achievement:** Dirac's equation for free electrons (and positrons)

$$E^2 = (mc^2)^2 + (pc)^2$$

$$E = \pm \sqrt{p^2 c^2 + m^2 c^4}$$

- ❖ **Interpretation of negative energies:**  
sea of electrons → holes in sea act as positively charged electrons  
→ confirmed by Anderson 1932

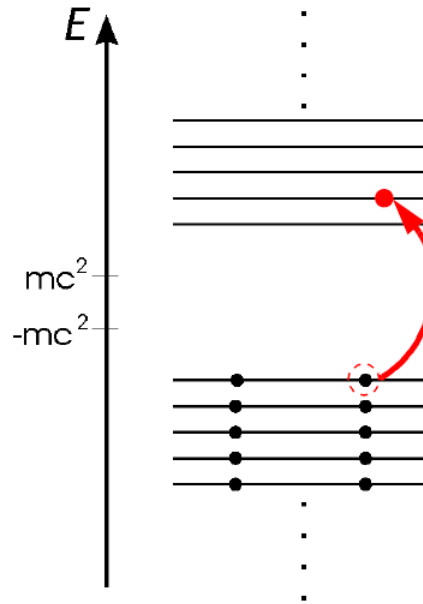


Paul Dirac  
1902-1984  
Nobel Prize 1933



Carl Anderson  
1905-1991  
Nobel Prize 1936

# Dirac's Picture of Vacuum



Fermions in Dirac's representation

The “hole” created by the appearance of the electron with a positive energy is interpreted as the presence of electron's *antiparticle* with the opposite charge.

“negative” energy  $\rightarrow$  **antiparticles**

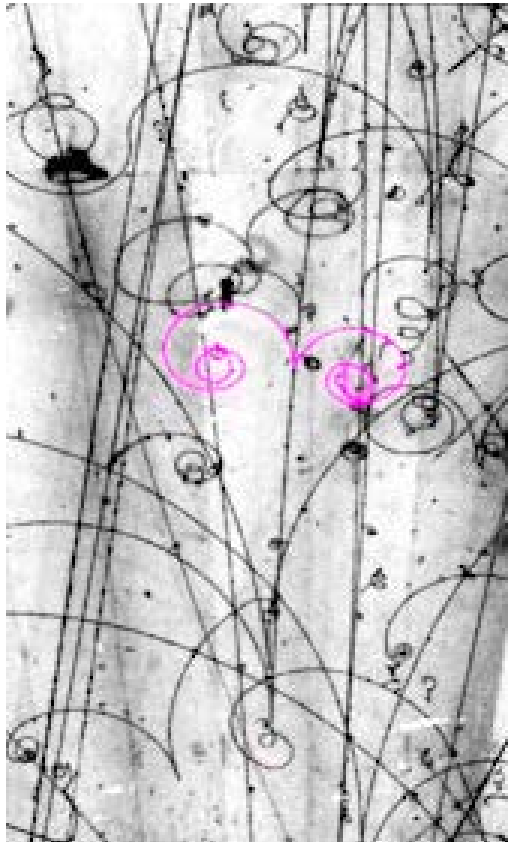
Later, Feynman – Stückelberg interpretation: antiparticles are particles travelling in reverse time direction

- Every charged particle has the antiparticle of the same mass and opposite charge.



# Antimatter

- ❖ Antiparticles look and behave just like their corresponding matter particles, except they have opposite charges.



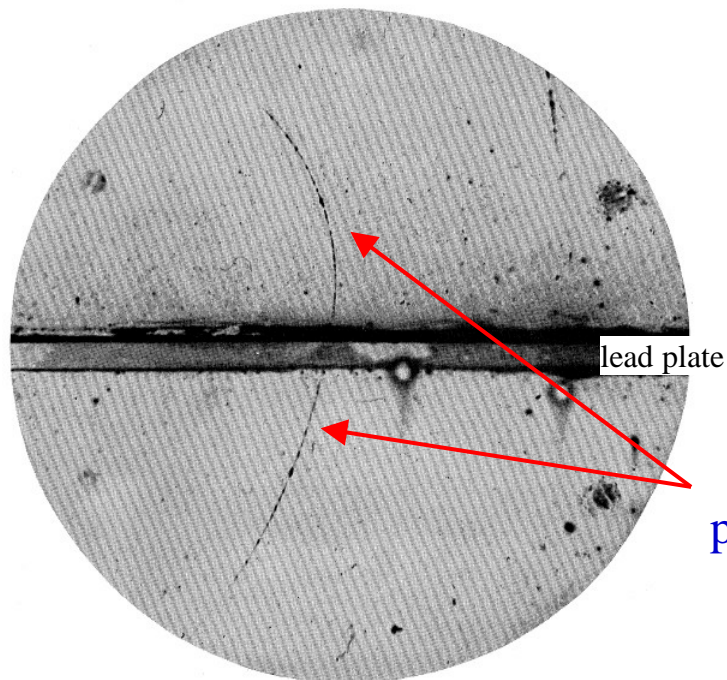
evidence for antimatter in this early **bubble chamber photo**.

The **magnetic field** in this chamber makes negative particles curl left and positive particles curl right.

The antielectron is called *positron* and is designated  $e^+$ .

# Discovery of the Positron

- ❖ First observed by D. Skobeltsyn in cosmic rays, using a Wilson cloud chamber (1929)
- ❖ C. D. Anderson: study of cosmic rays in a cloud chamber (1932)



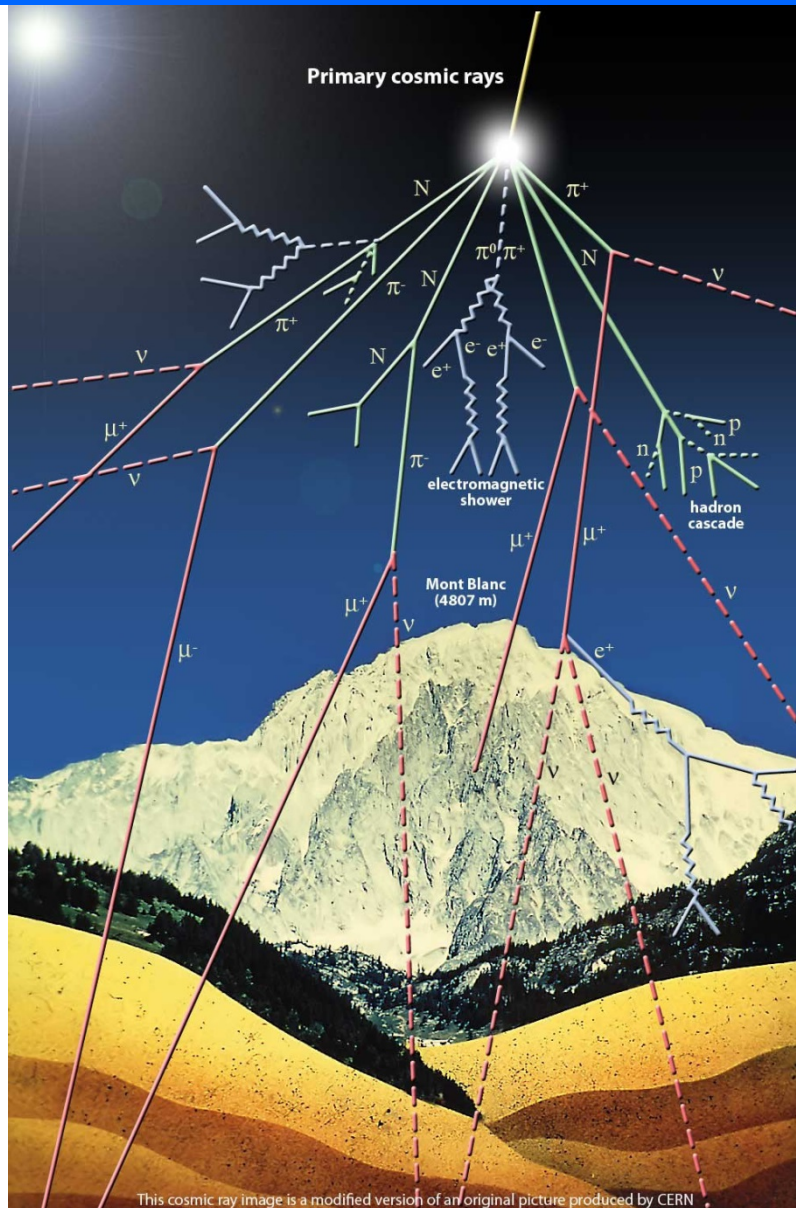
positrons are bent due to  
Lorentz force in magnetic field

proof of existence of antiparticles

positron track

FIG. 1. A 63 million volt positron ( $H\rho = 2.1 \times 10^6$  gauss-cm) passing through a 6 mm lead plate and emerging as a 23 million volt positron ( $H\rho = 7.5 \times 10^4$  gauss-cm). The length of this latter path is at least ten times greater than the possible length of a proton path of this curvature.

# Muon Discovery (1936)



Carl D. Anderson; Nobel Prize 1936

*Muon-Lepton* discovery in cosmic rays by

**Carl D. Anderson**

+

**Seth Nedermeyer**

mass:  $m_\mu \sim 100 \text{ MeV}$

- ❖ The muon is the first “exotic” particle
- ❖ It was believed to be the meson predicted by Yukawa

# Prediction of the Pion

H. Yukawa predicted 1935 mesons as carriers of the strong force

$$\text{meson mass} \sim \hbar c / (\text{range of force})$$

$$\text{range} \sim 1 \text{ fm} \rightarrow m_{\text{meson}} = 200 \text{ MeV}$$

$$\hbar c = 197.326 \text{ MeV fm}$$

note 1: *mesons are quark-antiquark  $q\bar{q}$*  states and not elementary

note 2: meson = “mean, intermediate state”



Hideki Yukawa  
1907-1981

The **range of forces** is related to the mass of exchange particle  $M$ . An amount of energy  $\Delta E = Mc^2$  borrowed for a time  $\Delta t$  is governed by the **Uncertainty Principle**:

$$\Delta E \cdot \Delta t \sim \hbar$$

The maximum distance the particle can travel is  $\Delta x = c \cdot \Delta t$ , where  $c$  is the velocity of light.

$$\Delta x = \hbar c / \Delta E$$

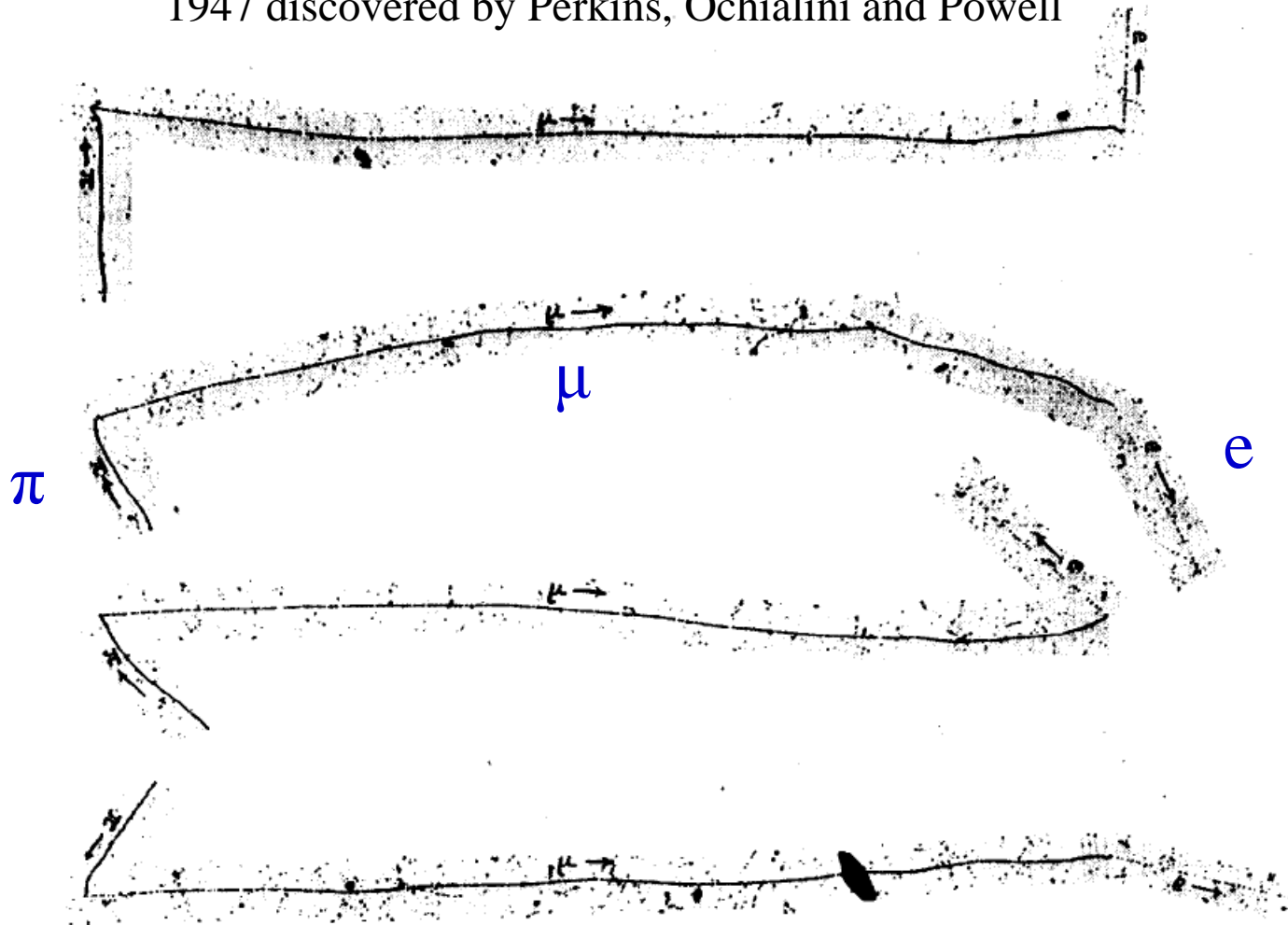
$$\Delta x = \hbar c / Mc^2$$

- ❖ The photon has  $M = 0$  → infinite range of EM force.
- ❖ W boson has a mass of  $80 \text{ GeV}/c^2$  → range of weak force  $\Delta x = 2 \cdot 10^{-3} \text{ fm}$

Nobel Prize 1949

# Discovery of the Pion

1947 discovered by Perkins, Ochialini and Powell



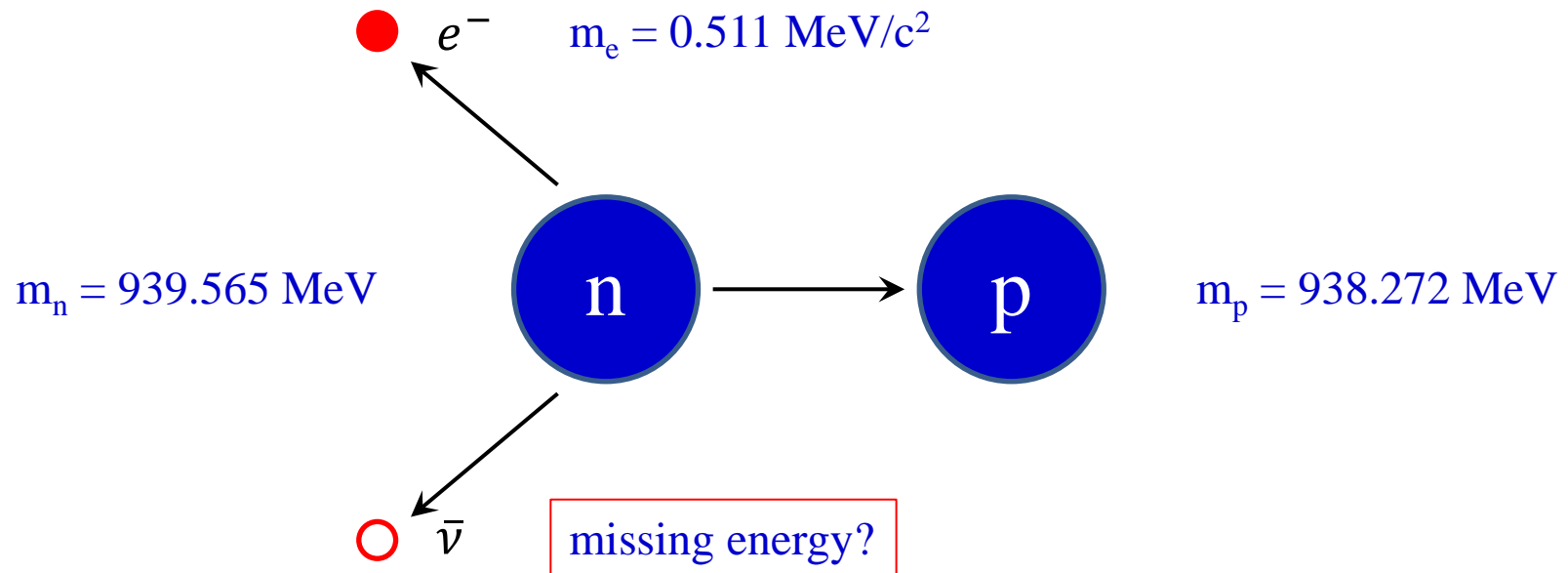
First observed pions: a  $\pi^+$  stops in the emulsion and decays to a  $\mu^+$  and  $\nu_\mu$ , followed by the decay of  $\mu^+$ .

Powell; Nobel Prize 1950



# Beta Decay

beta decay:  $n \rightarrow p + e^- + \bar{\nu}$

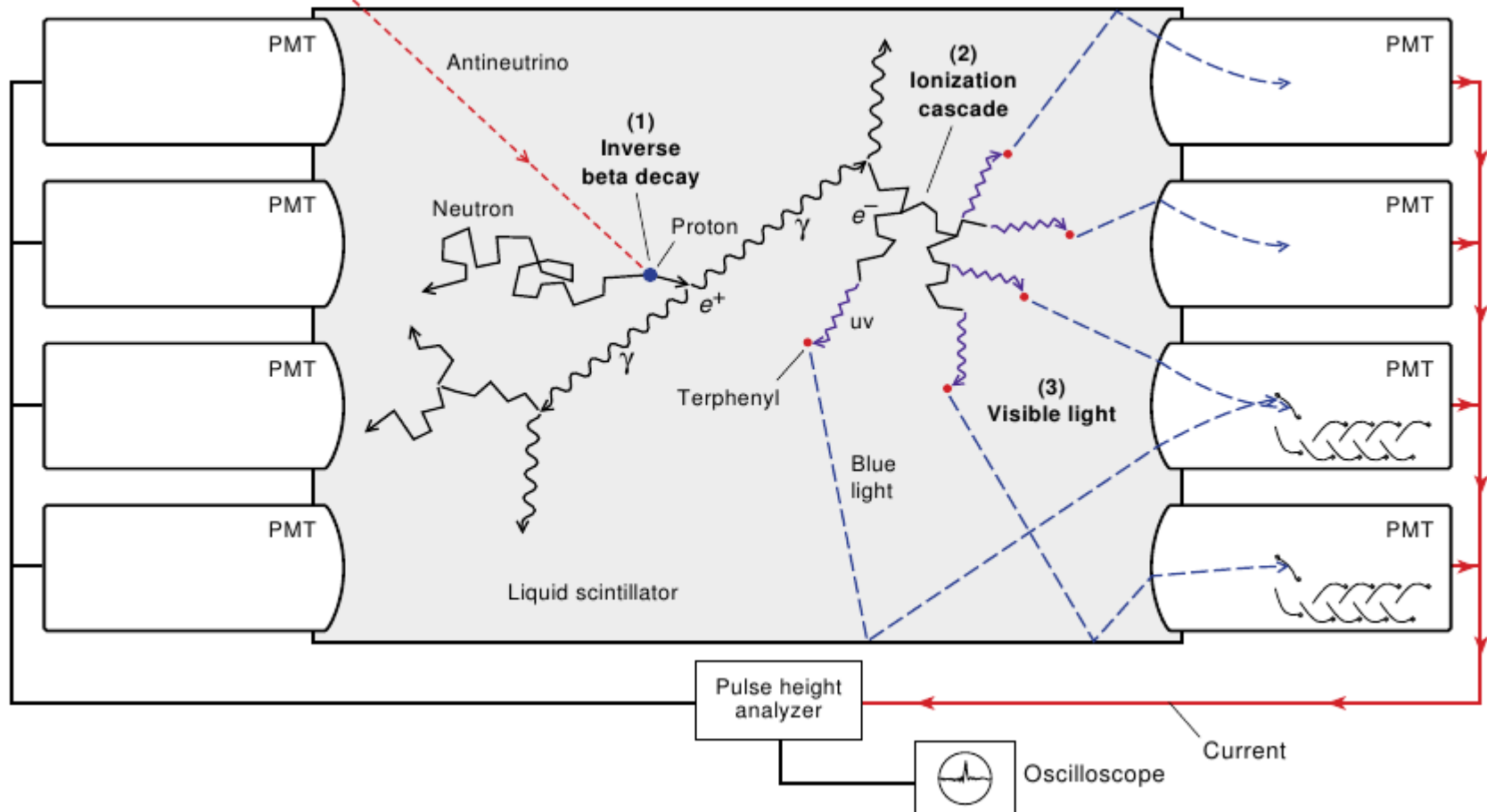


W. Pauli predicted in a letter to E. Fermi the existence of neutrinos (1930)

# Discovery of the Electron-Neutrino

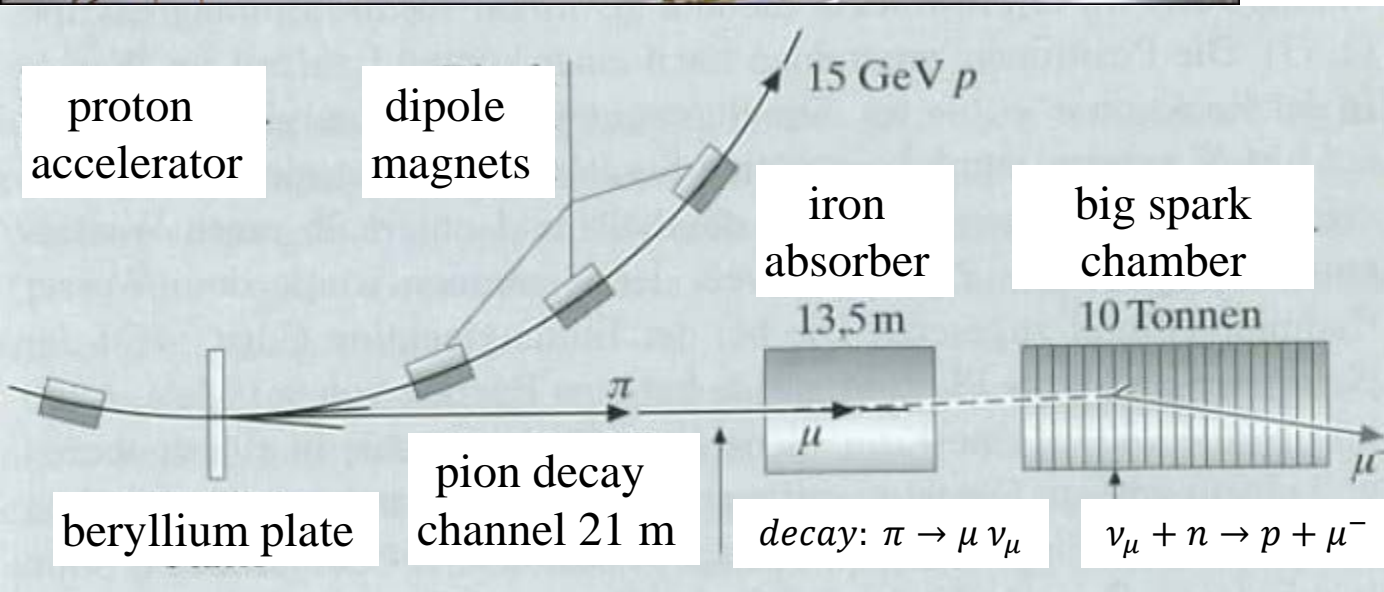
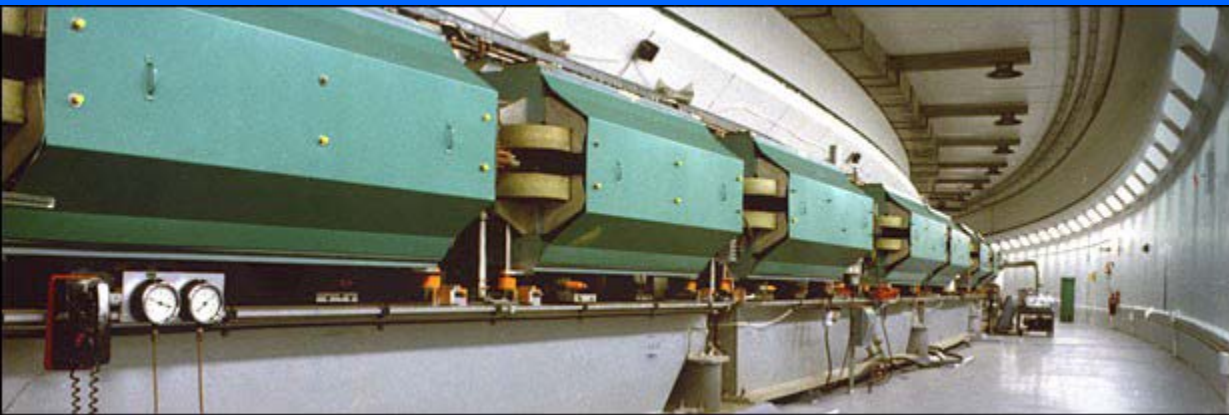
Discovery of the *electron neutrino* (Cowan, **Reines**, 1957)

Anti-Electron Neutrino



Nobel Prize 1995

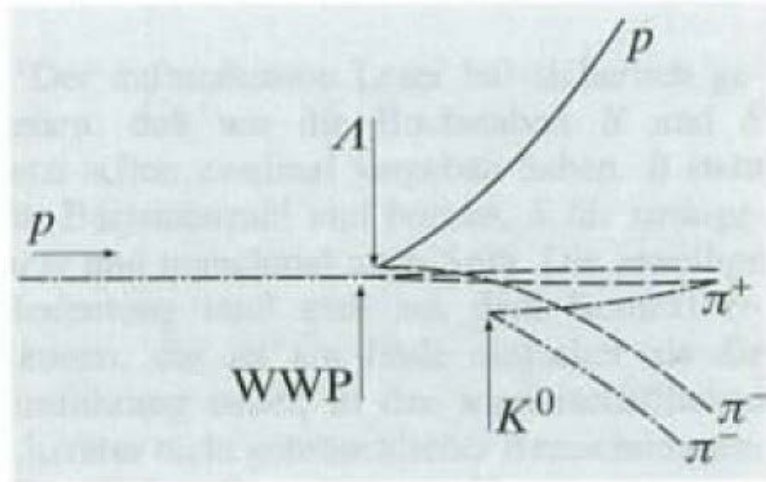
# Existence of Muon-Neutrinos



Discovery of the *muon neutrino* (Ledermann, Schwartz, Steinberger 1962)

Nobel Prize 1988

# Discovery of Strangeness Surprise!



“V-particles”

Production of *Kaons* and  $K^+(u\bar{s}), K^0(d\bar{s}/s\bar{d}), K^-(s\bar{u})$

*Lambda*-Baryons ( $uds$ ) in pp-collisions

Long lifetime!!!

neither electromagnetic nor strong force

(1950ties)

How does a bubble chamber work:

- It is filled with a liquid under pressure (hydrogen)
- Particles ionize the liquid along their passage
- When pressure drops, liquid boils preferentially along the ionization trails

# Discovery of Strangeness

## *Strange mesons and baryons*

were called so because, being produced in strong interactions, had quite long lifetimes and decayed weakly rather than strongly.

The most light particles containing s-quark are:

- ❖ mesons  $K^+$ ,  $K^-$  and  $K^0$ ,  $\bar{K}^0$ : “Kaons”, lifetime of  $K^+$  is  $1.2 \cdot 10^{-8}$  s
- ❖ baryon  $\Lambda$ , lifetime of  $2.6 \cdot 10^{-10}$  s

Principle decay modes of strange hadrons:

$$K^+ \rightarrow \mu^+ + \nu_\mu \quad (\text{B}=0.64)$$

$$K^+ \rightarrow \pi^+ + \pi^0 \quad (\text{B}=0.21)$$

$$\Lambda \rightarrow \pi^- + p \quad (\text{B}=0.64)$$

$$\Lambda \rightarrow \pi^0 + n \quad (\text{B}=0.36)$$

While the first decay in the list is clearly a weak one, decays of  $\Lambda$  can be very well described as strong ones, if not the long lifetime:  $(udd) \rightarrow (d\bar{u}) + (uud)$  must have a lifetime of order  $10^{-23}$  s, thus  $\Lambda$  can not be another sort of neutron...



# Status in 1962 – (50 years after Hess 'balloon trip)

known fermions

M. Gell-Mann and G. Zweig  
hypnotized existence of quarks:

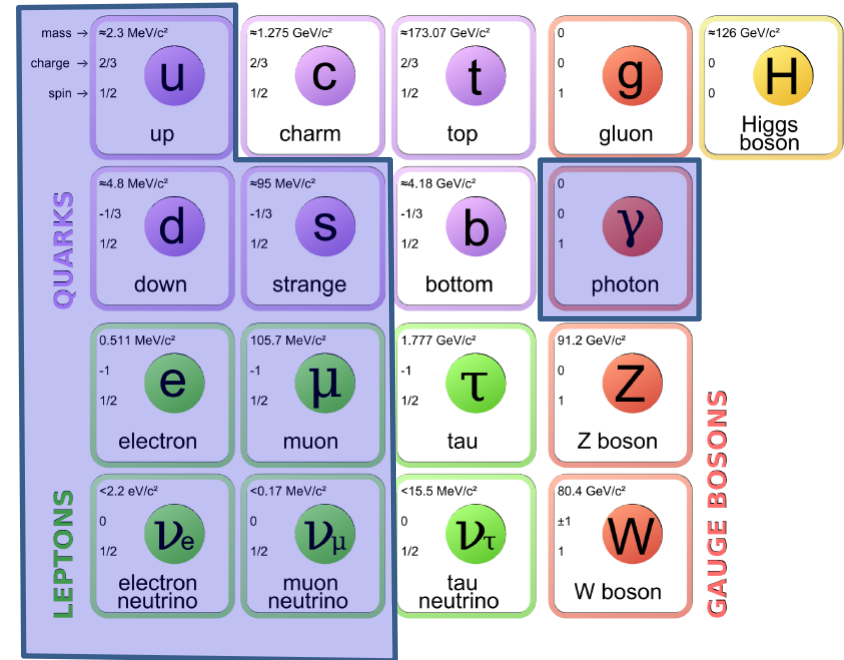
- up-quark
  - down-quark
  - strange-quark
- in the year 1964



Murray Gell-Mann  
1929-



George Zweig  
1937-

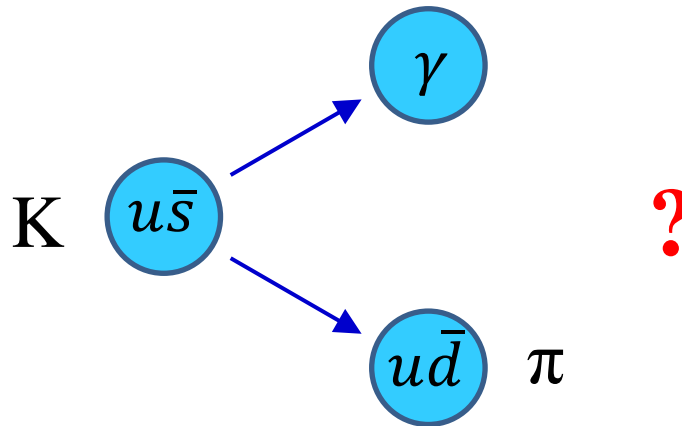


Nobel Prize 1969

# Prediction of the Charm-Quark

How explain non-observation of Flavor Changing Neutral Currents?

$$K \not\rightarrow \pi \gamma$$

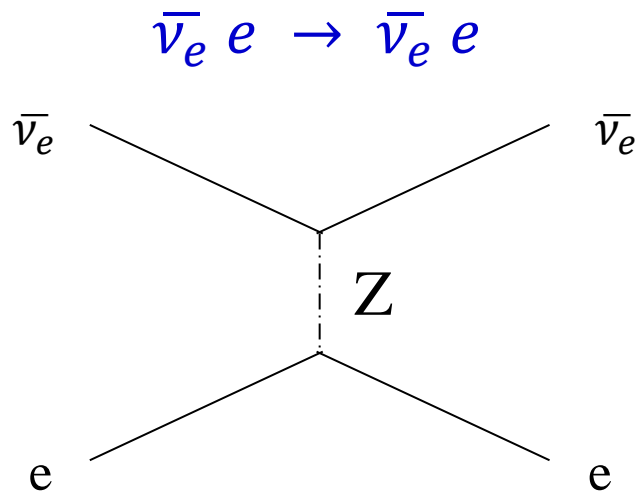


Such a **flavor changing** decay is forbidden if a **fourth quark** exists  
(Glashow, Iliopoulos, Maiani)

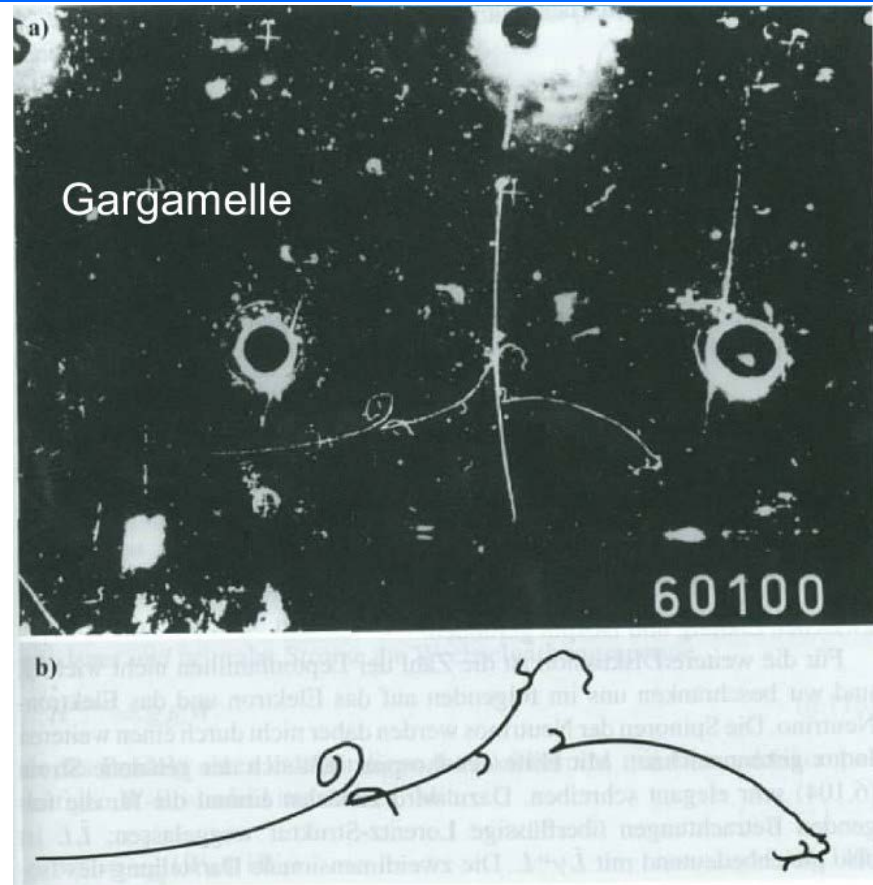
prediction

Note: Kaon and Pion have similar quantum numbers!

# Observation of Neutral Currents



Bubble Chamber Gargamelle (1974)



We find interactions without visible incoming or outgoing particle.

# Observation of the Charm-Quark

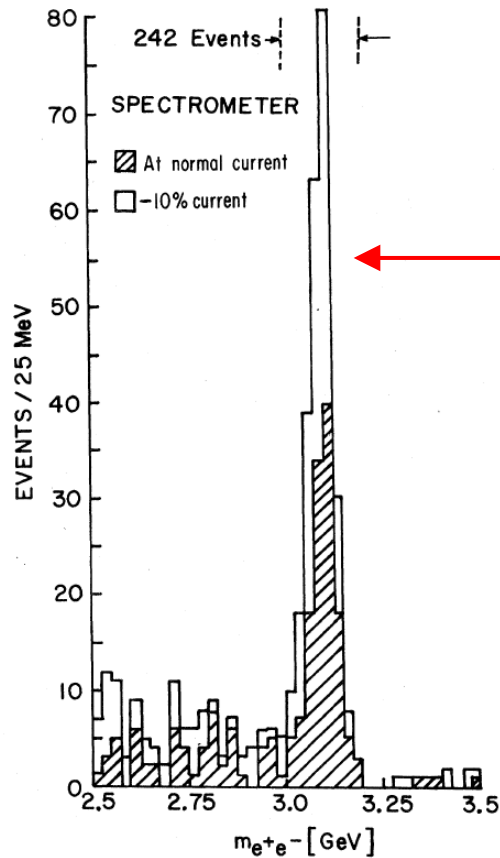


FIG. 2. Mass spectrum showing the existence of  $J$ . Results from two spectrometer settings are plotted showing that the peak is independent of spectrometer currents. The run at reduced current was taken two months later than the normal run.

$p \text{ Be} \rightarrow X \text{ J/Psi} \rightarrow X \text{ e}^+ \text{e}^-$   
BNL: **S. Ting** et al. (1974)

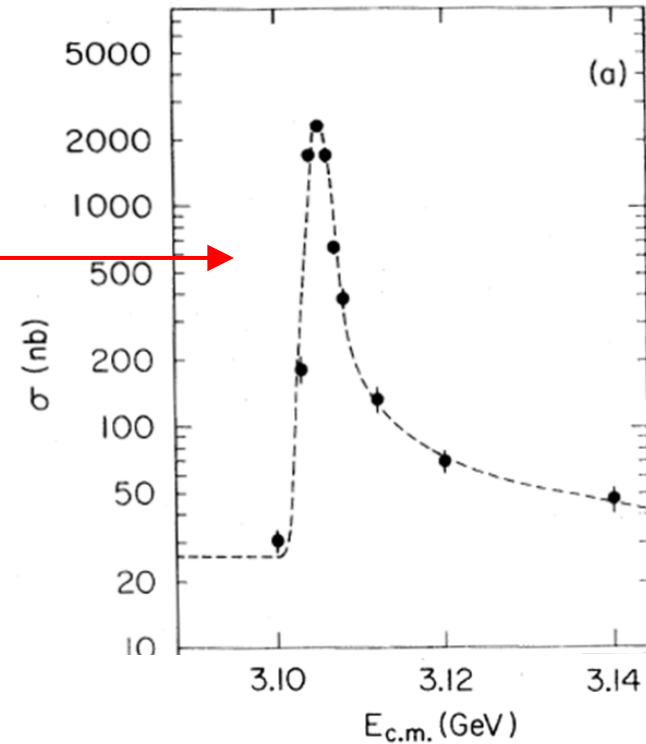


FIG. 1. Cross section versus energy for (a) multi-hadron final states, (b)  $e^+e^-$  final states, and (c)  $\mu^+\mu^-$ ,  $\pi^+\pi^-$ , and  $K^+K^-$  final states. The curve in (a) is the expected shape of a  $\delta$ -function resonance folded with the Gaussian energy spread of the beams and including radiative processes. The cross sections shown in (b) and (c) are integrated over the detector acceptance. The total hadron cross section, (a), has been corrected for detection efficiency.

$e^+e^- \rightarrow \text{J/Psi} \rightarrow e^+e^- (\mu^+\mu^-), (\pi^+\pi^-)$   
SLAC: **B. Richter** et al. (1974)

Nobel Prize 1976

# Discovery of Charm-Quark

- ❖ Bubble chambers were great tools of particle discovery, providing physicists with numerous hadrons, all of them fitting  $u-d-s$  quark scheme until 1974.
- ❖ In 1974, a new particle was discovered, which demanded a new flavour to be introduced. Since it was detected simultaneously by two rival groups in Brookhaven (BNL) and Stanford (SLAC), it received a double name:

$$J/\Psi (3097) = c\bar{c}$$

The new quark was called “*charmed*”, and the corresponding quark number is *charm*,  $C$ . Since  $J/\Psi$  itself has  $C = 0$ , it is said to contain “hidden charm”.

Shortly after that particles with “naked charm” were discovered as well:

$$D^+(1869) = c\bar{d}, \quad D^0(1865) = c\bar{u}$$

$$D^-(1869) = d\bar{c}, \quad D^0(1865) = cu$$

$$\Lambda_c^+(2285) = udc$$

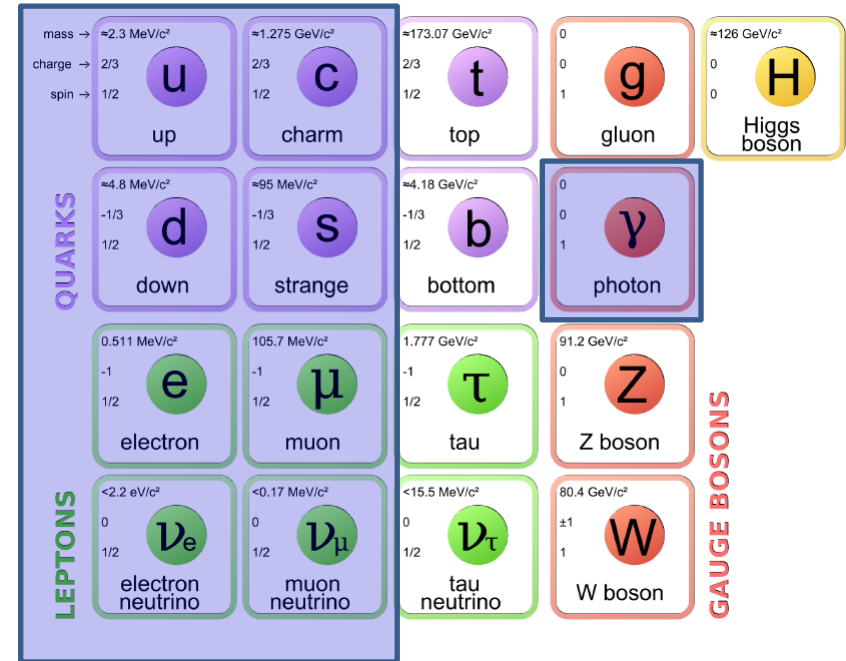


# Status in 1975

known fermions



no anti-matter in universe!



Charge and Parity Conservation must be broken!  
(2<sup>nd</sup> A. Sakharovs condition)

→ **third quark family** could explain matter – anti-matter asymmetry  
Nambu, Kobayashi, Maskawa

Note: particle physics and evolution of universe are closely related!

Nobel Prize 2008

# Discovery of the Third Family

**Tau Lepton** discovered by MARK1 at SPEAR (1974-1976)

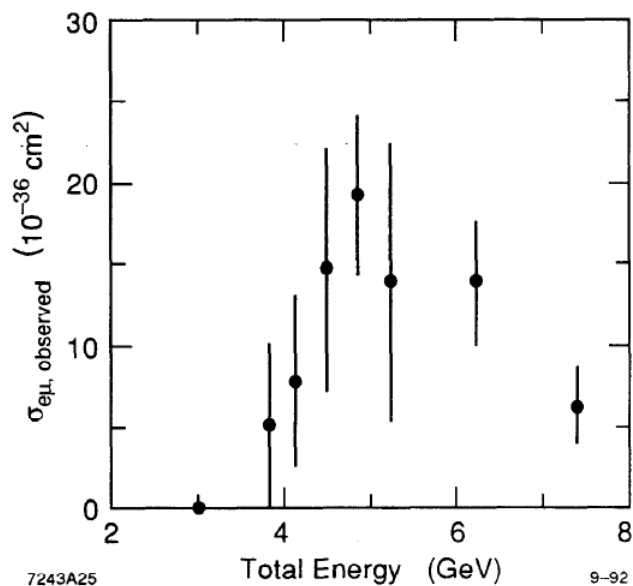


Fig. 4. From Perl *et al.* (1975): the observed cross section for the signature  $e\mu$  events from the Mark I experiment at SPEAR. This observed cross section is not corrected for acceptance. There are 86 events with a calculated background of 22 events.

$$e^+e^- \rightarrow \tau^+\tau^-$$

$$\tau \rightarrow e \nu \nu$$

$$\tau \rightarrow \mu \nu \nu$$

$$m_\tau = 1.777 \text{ GeV}$$

discovery of *tau-lepton*

(Reines *et al.*, M. Perl Nobel Prize 1995)

Discovery of the *Bottom-Quark* (1977) by M. Ledermann *et al.* Nobel Prize 1988



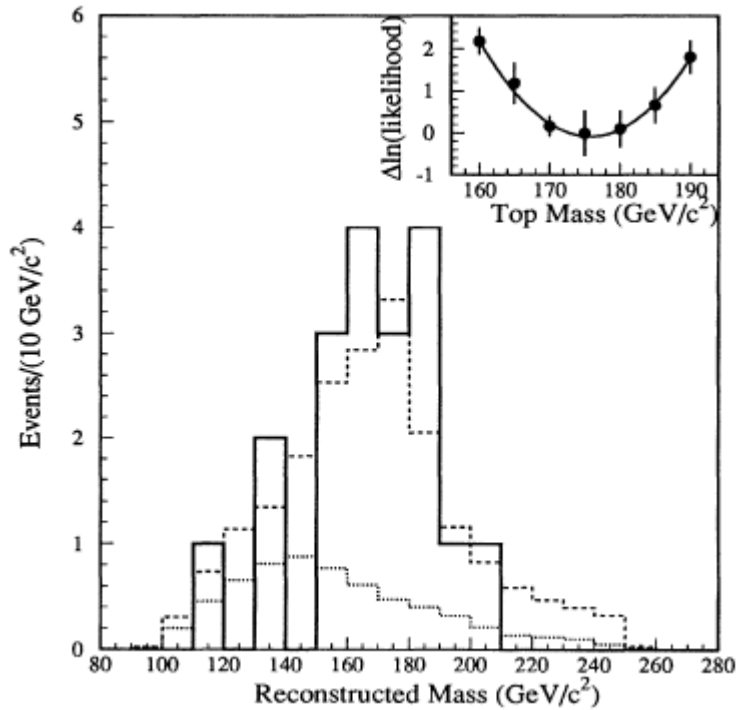
# Top-Quark Discovery at Tevatron



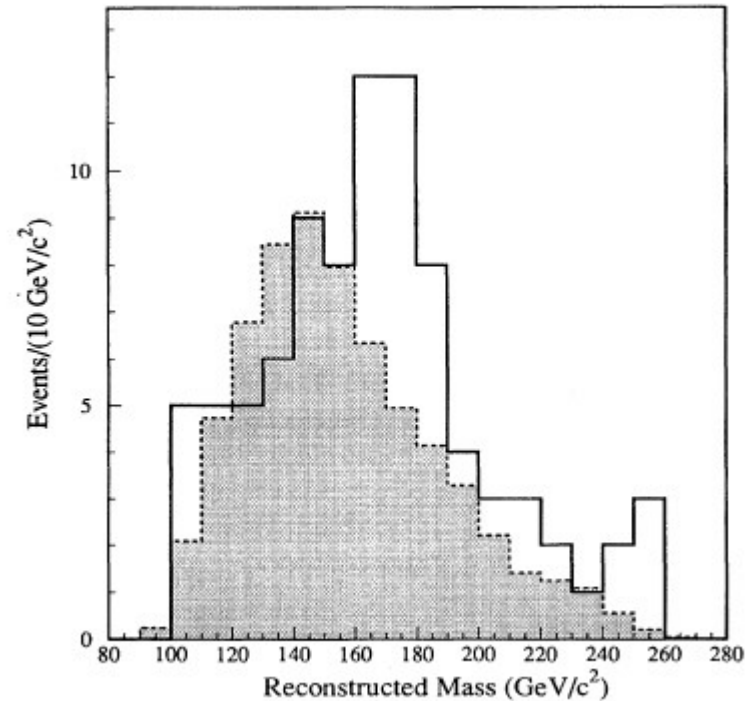
Fermilab

Tevatron  
 $p\bar{p}$  collider  $s^{1/2} = 2 \text{ TeV}$

# Top-Quark Discovery in 1995



CDF Collaboration



C0 Collaboration

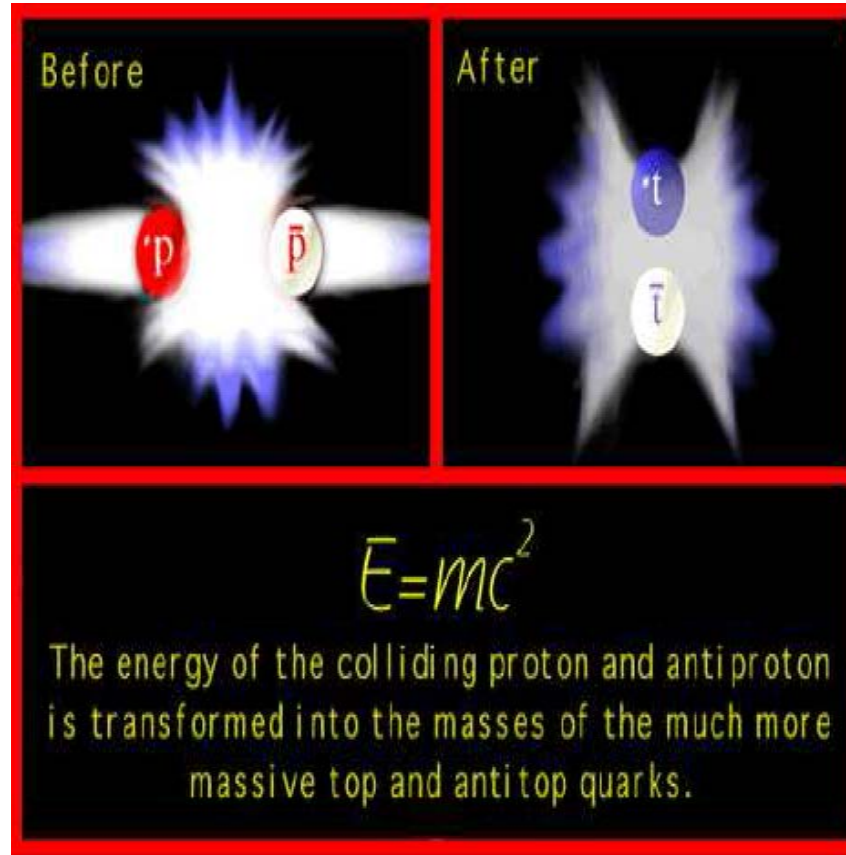
**Top-Quark mass ~ 175 GeV**

Top production:  $p\bar{p} \rightarrow t\bar{t} X$

Top decay:  $t \rightarrow W b$  and  $W \rightarrow q\bar{q}$



# Top-Quark Discovery in 1995



**Top-Quark** mass  $\sim 175$  GeV

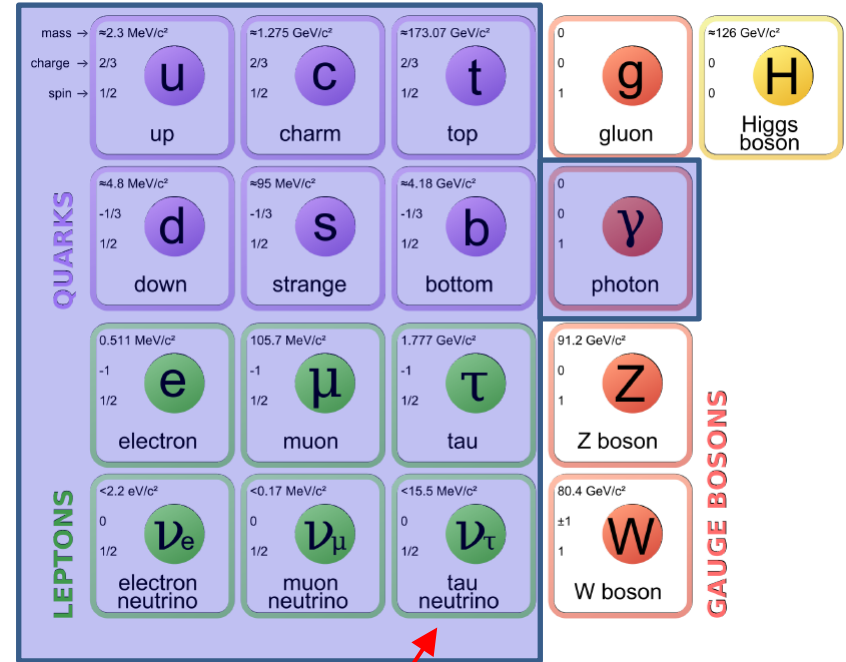
Top production:  $p\bar{p} \rightarrow t\bar{t} X$

Top decay:  $t \rightarrow W b$  and  $W \rightarrow q\bar{q}$



# Status in 2000 (~90 years after Hess balloon trip)

all fermions discovered



*tau-neutrino* discovered  
in 2000 by Donut experiment

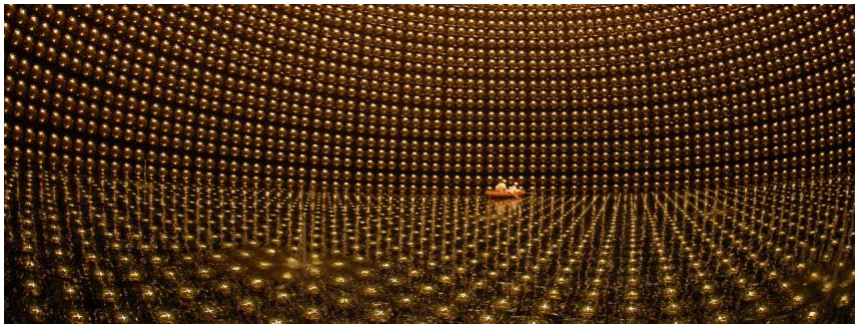
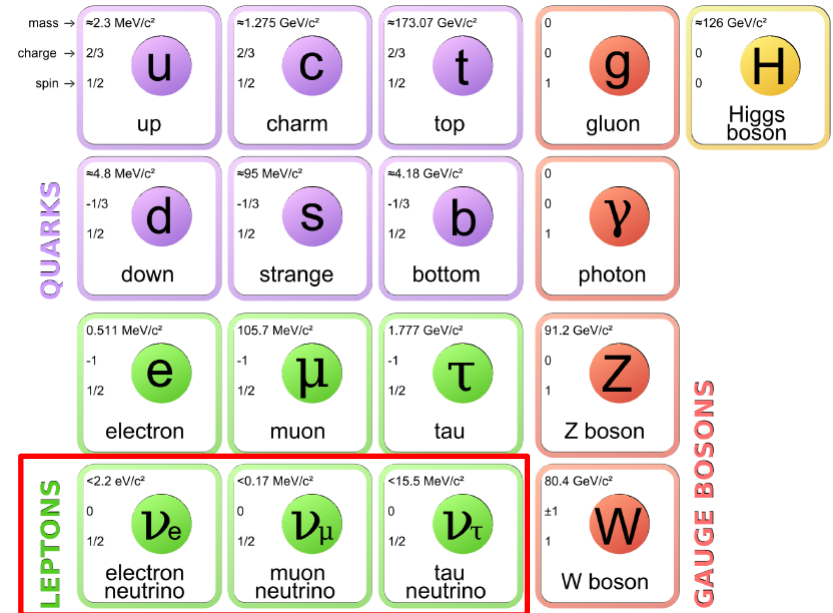
# Status in 2002 (~90 years after Hess balloon trip)

Nobel Prize 2015

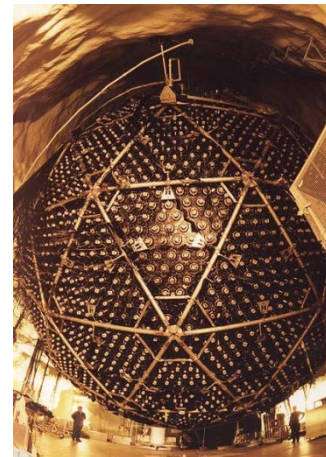


Takaaki Kajita + Arthur B. McDonald

neutrinos have mass!



Super-Kamiokande (1998)



SNO experiment (2001/2)

# Discovery of Vector-Bosons

## Photons:

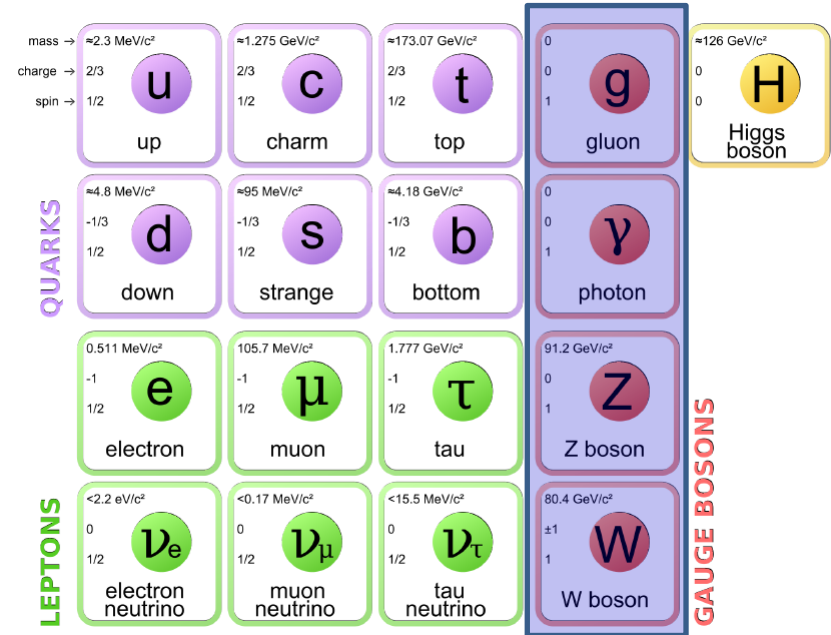
- X-rays (Röntgen 1895) → Nobel Prize 1901
- photo-electric effect (Einstein 1905) → Nobel Prize 1921

## Gluons:

- 8 carriers of strong force
- predicted by Gell-Mann et al.
- experimental discovery as jets of hadrons (e.g. PETRA)

## $W^\pm$ bosons:

- carriers of weak interactions
- responsible for weak decays
- building blocks of “**Standard Model**”  
Glashow, Salam, Weinberg → Nobel Prize 1979
- W-boson discovered at CERN  $Sp\bar{p}S$  in 1983  
C. Rubbia and S. van der Meer → Nobel Prize 1984

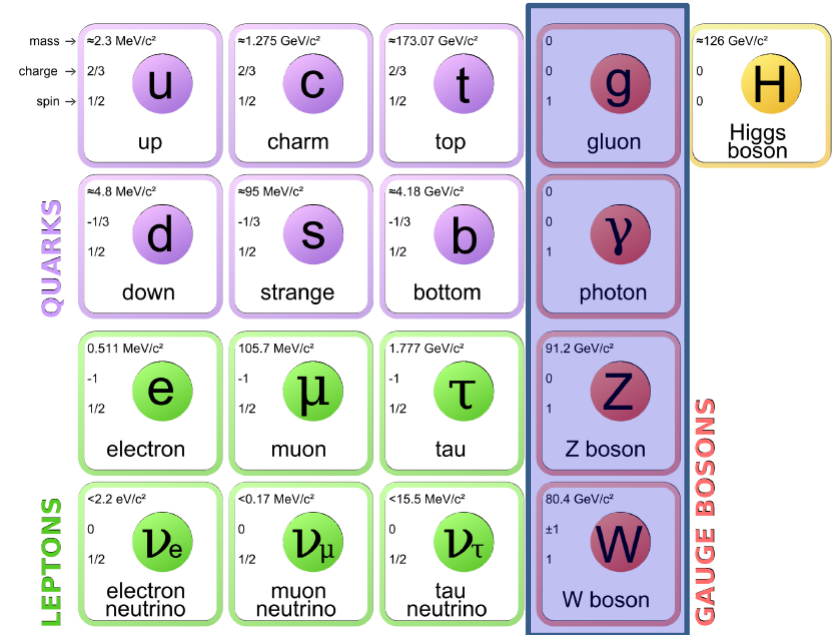


Abdus Salam, Steven Weinberg, Sheldon L. Glashow

# Discovery of Vector Bosons

## Z-boson:

- indirectly observed at Gargamelle
- discovered at CERN  $Spp\bar{S}$  in 1983  
C. Rubbia and S. van der Meer → Noble Prize 1984
- precision studies at LEP accelerator



Large Electron Positron (LEP) collider



Carlos Rubbia

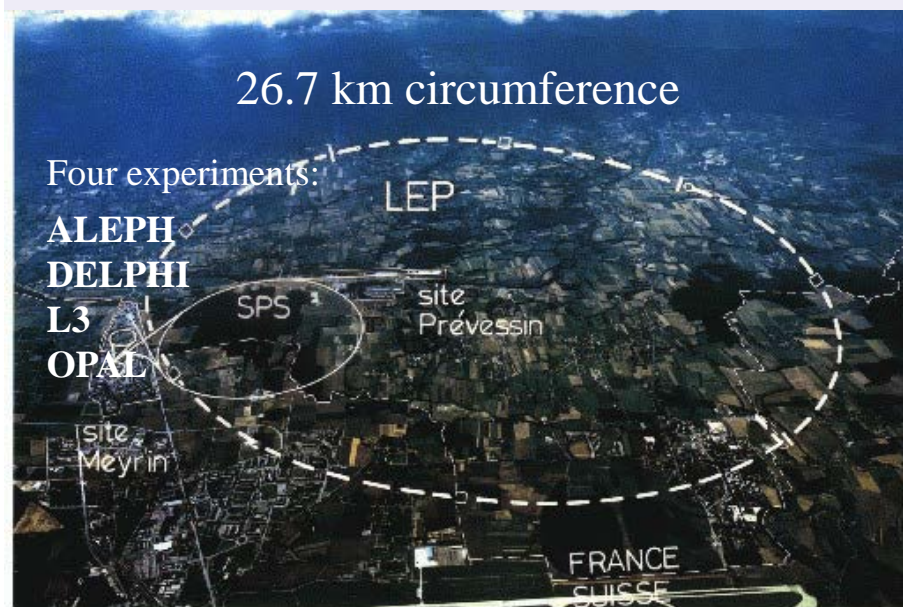
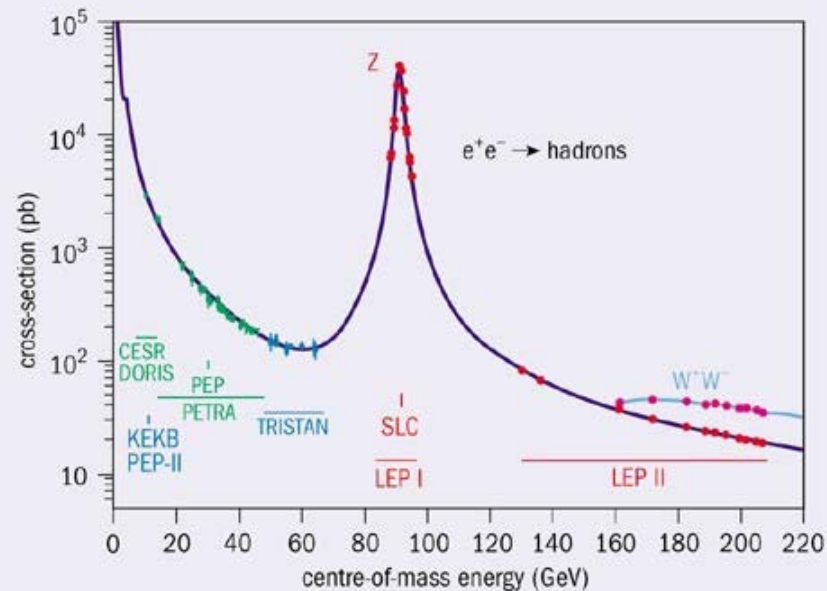
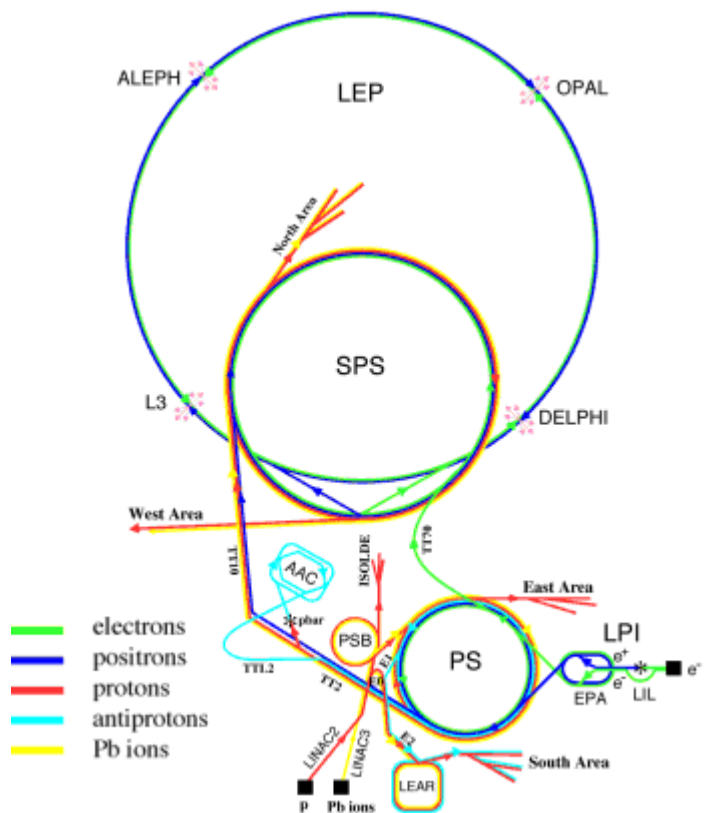


Simon van der Meer



# Large Electron-Positron Collider at CERN

1989-2000:  $S^{1/2} = 90\text{-}200\text{ GeV}$



# Large Hadron Collider at CERN



LHC ( $p \rightarrow \leftarrow p$ )  
 $s^{1/2} \leq 14 \text{ TeV}$

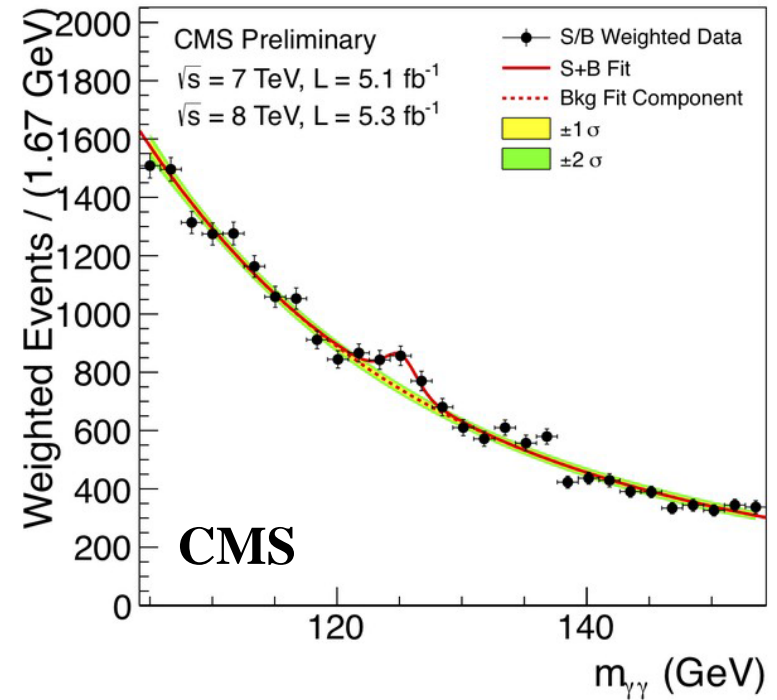
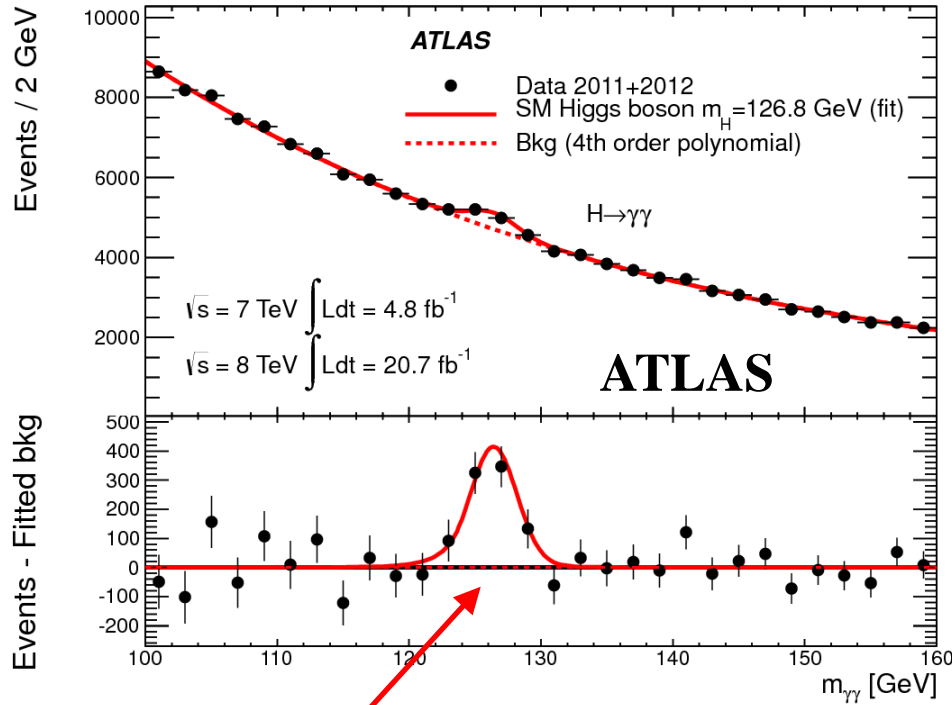
26.7 km circumference

- 40000 tons of cold mass spread over 27 km
- 10000 tons of Liquid Nitrogen (at  $T = 80 \text{ K}$ )
- 60 tons of Liquid Helium (cools ring to final  $1.9 \text{ K}$ )





# The Higgs-Boson



**$M_H = 125$  GeV**



Nobel Prize 2013  
 F. Englert  
 P. Higgs



# CMS - Collaboration

