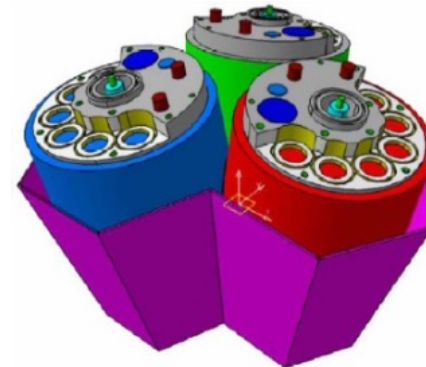
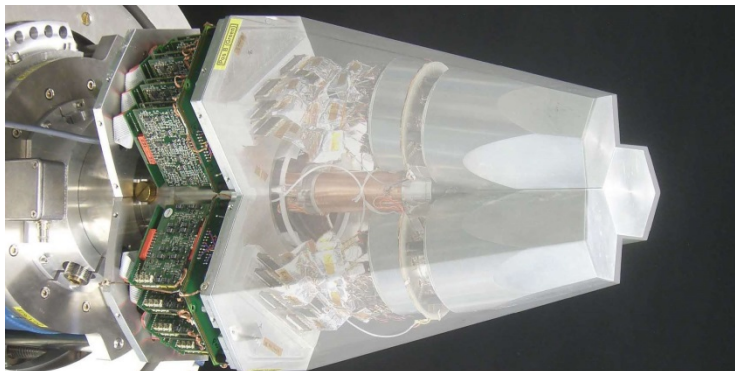
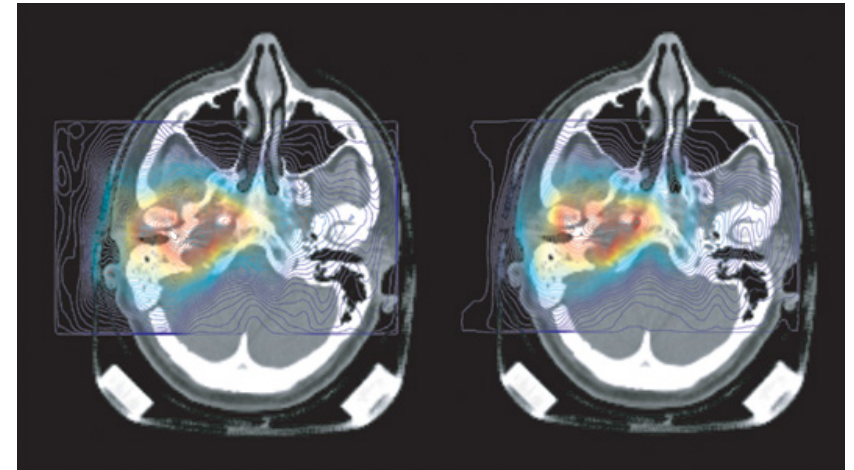
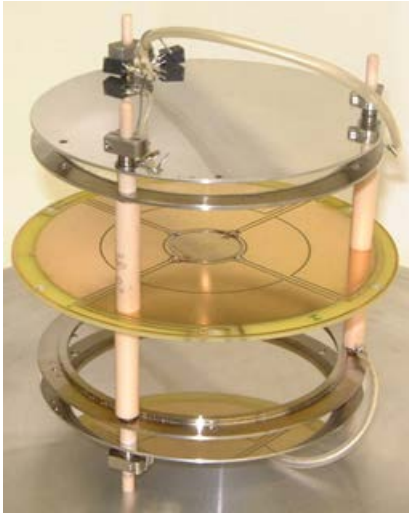


# Particle and Radiation Detectors: Advances & Applications

Lecture: Hans-Jürgen Wollersheim

e-mail: [h.j.wollersheim@gsi.de](mailto:h.j.wollersheim@gsi.de)



# Tentative outline of detector lecture

## ❖ Charged particle interaction in matter

relevant formulas

Bethe-Bloch formula

interaction of  $\beta$ -particles in matter

Bremsstrahlung

Synchrotron radiation

## ❖ Interaction of gamma-rays in matter

photoelectric effect

Compton scattering

$e^+e^-$  pair production

## ❖ Semiconductor detectors

bond model for semiconductors

doping: n- and p-type silicon

Silicon surface barrier detector

micro-strip detector

Germanium detector

## ❖ Segmented detectors

Doppler effect

Euroball, Miniball, AGATA

## ❖ Electronics

signal processing

constant fraction discriminator

coincidences

## ❖ Gas detectors

ionization detectors

ionization chamber

proportional counter

multi-wire proportional chamber

time projection chamber

Geiger-Müller counter

## ❖ Scintillation detectors

organic and inorganic scintillators

Liouville's theorem

photomultiplier tube

neutron detector

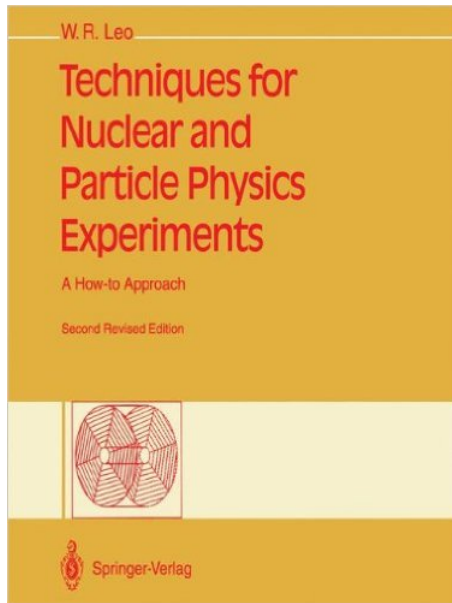
## ❖ Neutral particles

neutrons

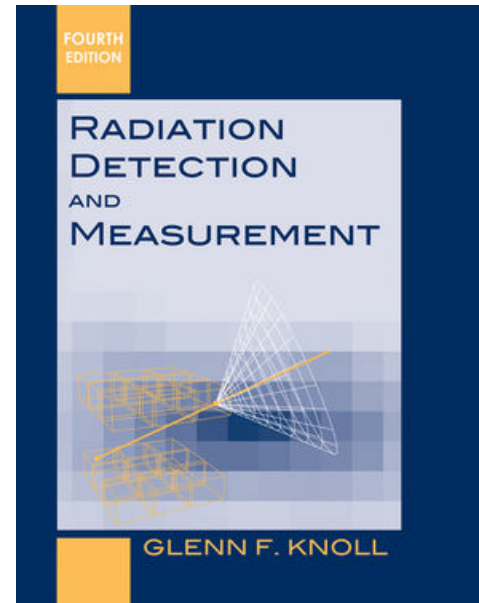
neutrino

## ❖ Imaging

# Literature



❖ Recommended Textbook

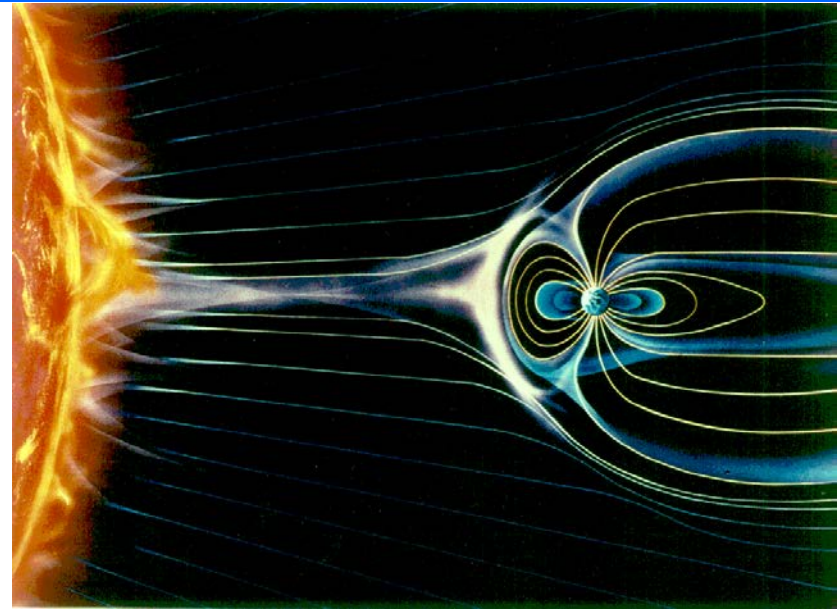


❖ Recommended Textbook

# Particle Interaction with Matter



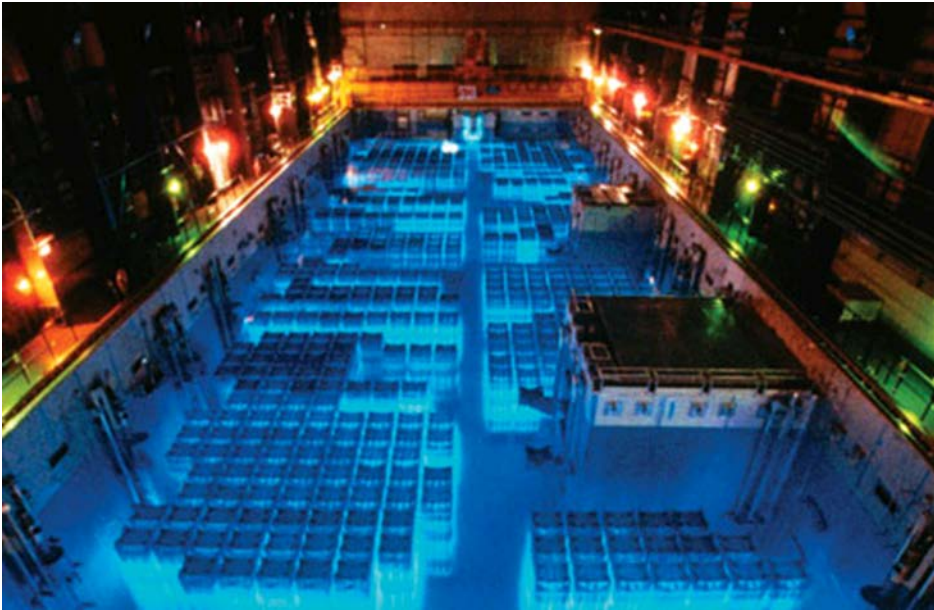
Aurora Borealis



Ionization



# Particle Interaction with Matter



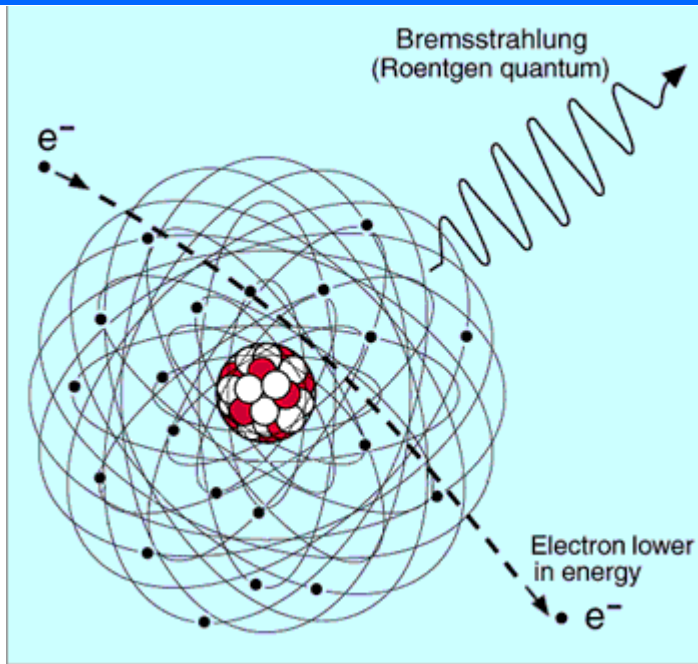
Characteristic glow from a reactor

Cherenkov Light



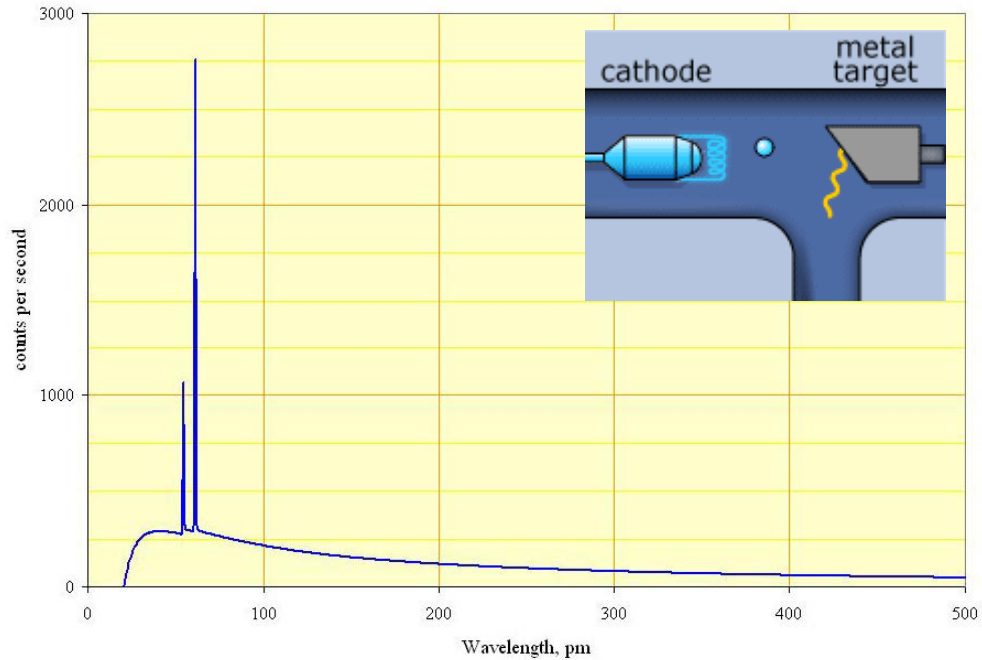
Cherenkov radiation is an effect similar to sonic booms when the plane exceeds the velocity of sound

# Particle Interaction with Matter



Bremsstrahlung or 'braking radiation'

**Bremsstrahlung**



Spectrum of the X-rays emitted by an X-ray tube with a rhodium target, operated at 60 kV. The continuous curve is due to bremsstrahlung, and the spikes are characteristic K lines for rhodium.

# Interaction of gamma rays with matter

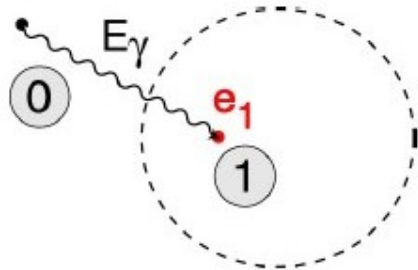
~ 100 keV

~1 MeV

~ 10 MeV

$\gamma$ -ray energy

## Photoelectric

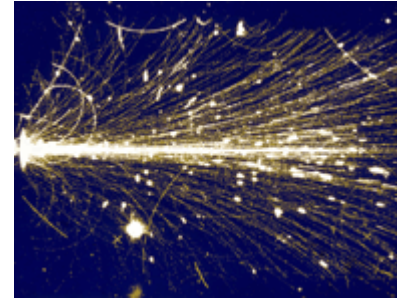


## Isolated hits

Probability of  
interaction depth

# Energy loss – dE/dx

Particles interact differently with matter. Important for detectors is the energy loss per path length. The total energy loss per path length is the sum of all contributions.



$$-\left(\frac{dE}{dx}\right)_{tot} = -\left(\frac{dE}{dx}\right)_{coll} - \left(\frac{dE}{dx}\right)_{rad} - \left(\frac{dE}{dx}\right)_{photoeff} - \left(\frac{dE}{dx}\right)_{compton} - \left(\frac{dE}{dx}\right)_{pair} - \left(\frac{dE}{dx}\right)_{hadron} \dots$$

Depending on the particle type, the particle energy and the material some processes dominate, other do not occur. For instance only charged particles will interact with electrons of atoms and produce ionization, etc.



# Measurement Principles

A measurement requires an interaction of the particle with the material of the detector. The interaction provokes two effects:

- 1<sup>st</sup>      **Creation of a detectable signal**, e.g.  
ionization → charges  
excitation → scintillation  
excitation of phonons → heat
  
- 2<sup>nd</sup>      **Alternation of the particles properties**, e.g.  
energy loss  
change of trajectory due to scattering  
absorption

unwanted side effects. They need to be as small as possible and well understood.

# Measurement Principles

*A particle detector is an instrument to measure one or more properties of a particle ...*

## Properties of a particle

- position and direction
- momentum
- energy
- mass
- velocity
- transition radiation
- spin, lifetime

$x, \vec{x}$

$|\vec{p}|$

$E$

$m$

$\beta$

$\gamma$

## Type of detection principle:

position and tracking

tracking in a magnetic field

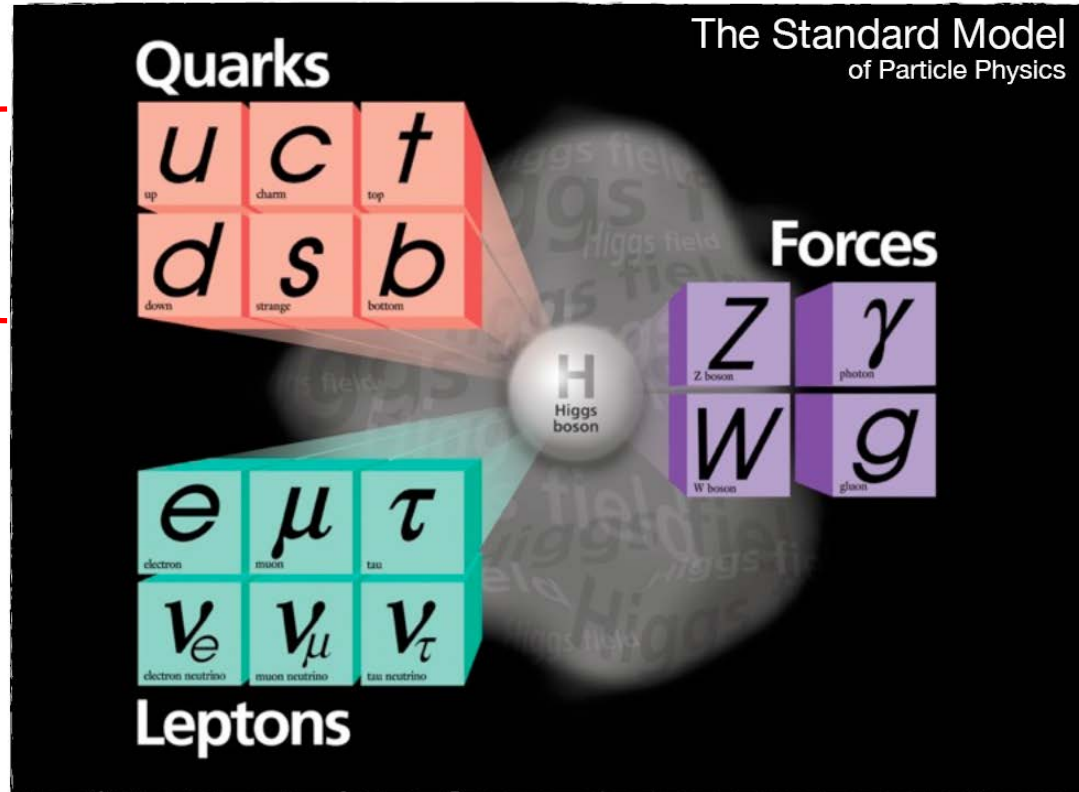
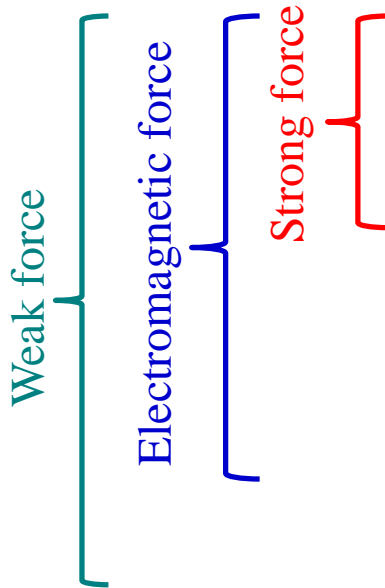
calorimetry

spectroscopy and PID

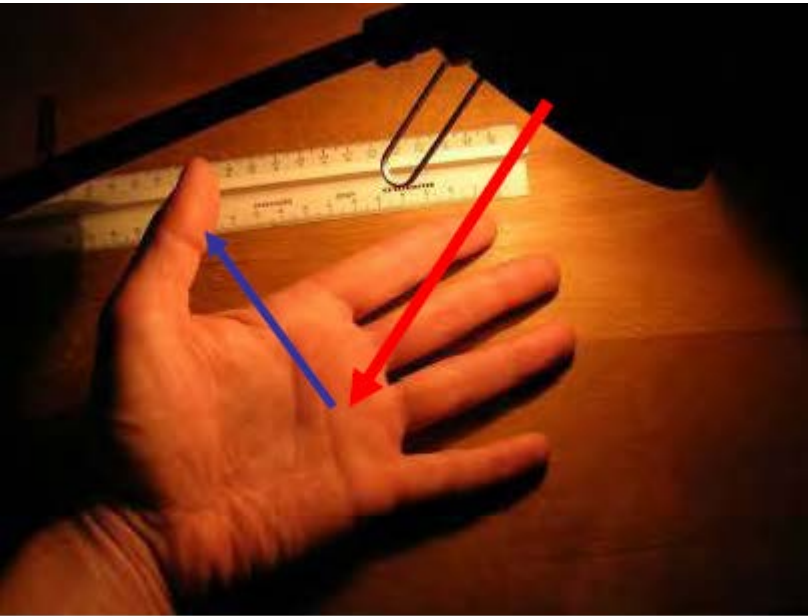
Cherenkov radiation or time of flight

TRD

# What is a Particle?



# How do we see Particles?



A light bulb shines on a hand and the different reflections make the fine structure visible.

With a magnifying glass or microscope more details can be seen, but there is a fundamental limit:

The wavelength of the light (1/1000 mm) determines the size of the resolvable objects.

available wavelength

→ electromagnetic waves  $E = \frac{hc}{\lambda}$

LW	3000 m	
MW	300 m	
KW	30 m	
UKW	3 m	
GPS	0.3 m	
Infrared	$10^{-6}$ m	
light	$5 \cdot 10^{-7}$ m	2 eV
UV	$10^{-7}$ m	10 eV
X-ray	$10^{-10}$ m	$10^4$ eV
$\gamma$ -ray	$10^{-12}$ m	$10^6$ eV

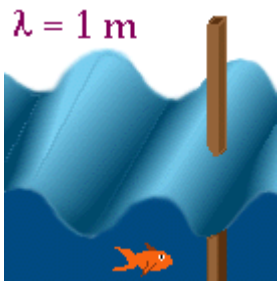
light bulb

→ **accelerator**

magnifying glass or microscope

→ **detector**

# Energy, Wavelength and Resolution



wavelength versus resolution

Small objects (smaller than  $\lambda$ ) do not disturb the wave

→ small object is not visible

Large objects disturb the wave

→ large object is visible

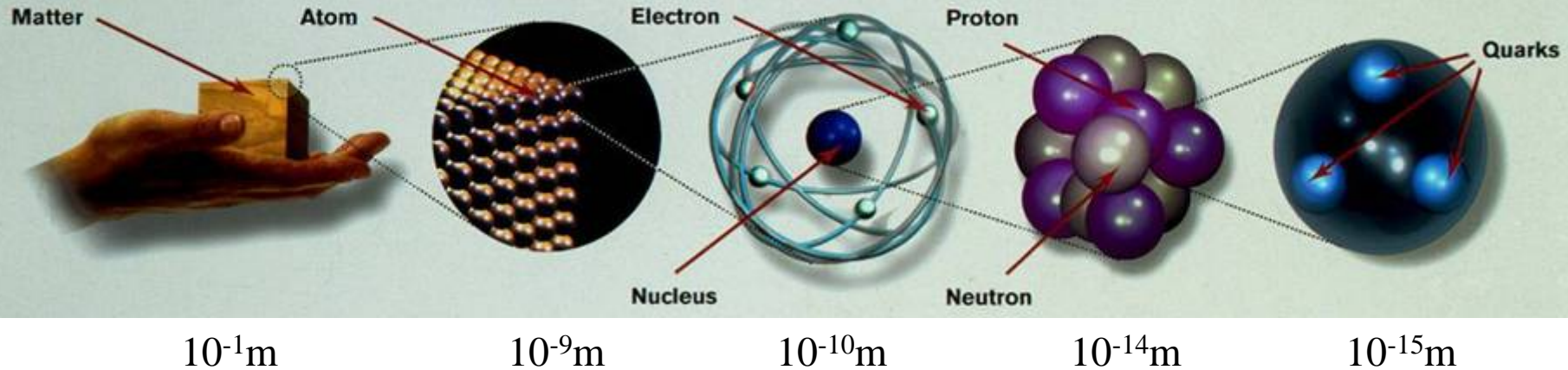
❖ all particles have wave properties:

$$\lambda = \frac{h}{p} = \frac{hc}{\sqrt{E_{kin} \cdot (E_{kin} + 2m_0c^2)}}$$

de Broglie wavelength



Louis de Broglie



$$h \cdot c = 1239.84 \text{ [MeV fm]}$$

# Detection and Identification of Particles

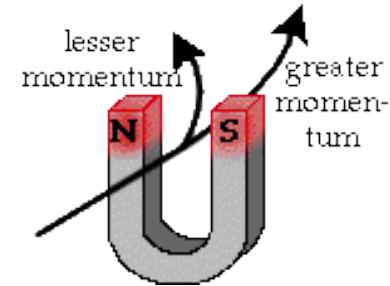
- ❖ **Detection** = particle counting (is there a particle?)
- ❖ **Identification** = measurement of **mass** and **charge** of the particle  
(most elementary particles have  $Ze = \pm 1$ )

❖ **How:**

- charged particles are deflected by B fields such that:

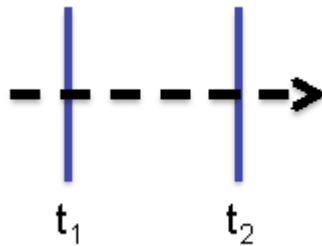


$$\rho = \frac{p}{ZeB} \propto \frac{p}{Z} = \frac{\gamma m_0 \beta c}{Z}$$



$p$  = *particle momentum*  
 $m_0$  = *rest mass*  
 $\beta c$  = *particle velocity*

- **particle velocity** measured with time-of-flight (ToF) method



$$\beta \propto \frac{1}{\Delta t}$$

- ❖ ToF for known distance
- ❖ Ionization  $-\frac{dE}{dx} = f(\beta)$
- ❖ Cherenkov radiation
- ❖ Transition radiation

# Detection and Identification of Particles

- ❖ **Detection** = particle counting (is there a particle?)
- ❖ **Identification** = measurement of **mass** and **charge** of the particle  
(most elementary particles have  $Ze = \pm 1$ )

- ❖ **How:**

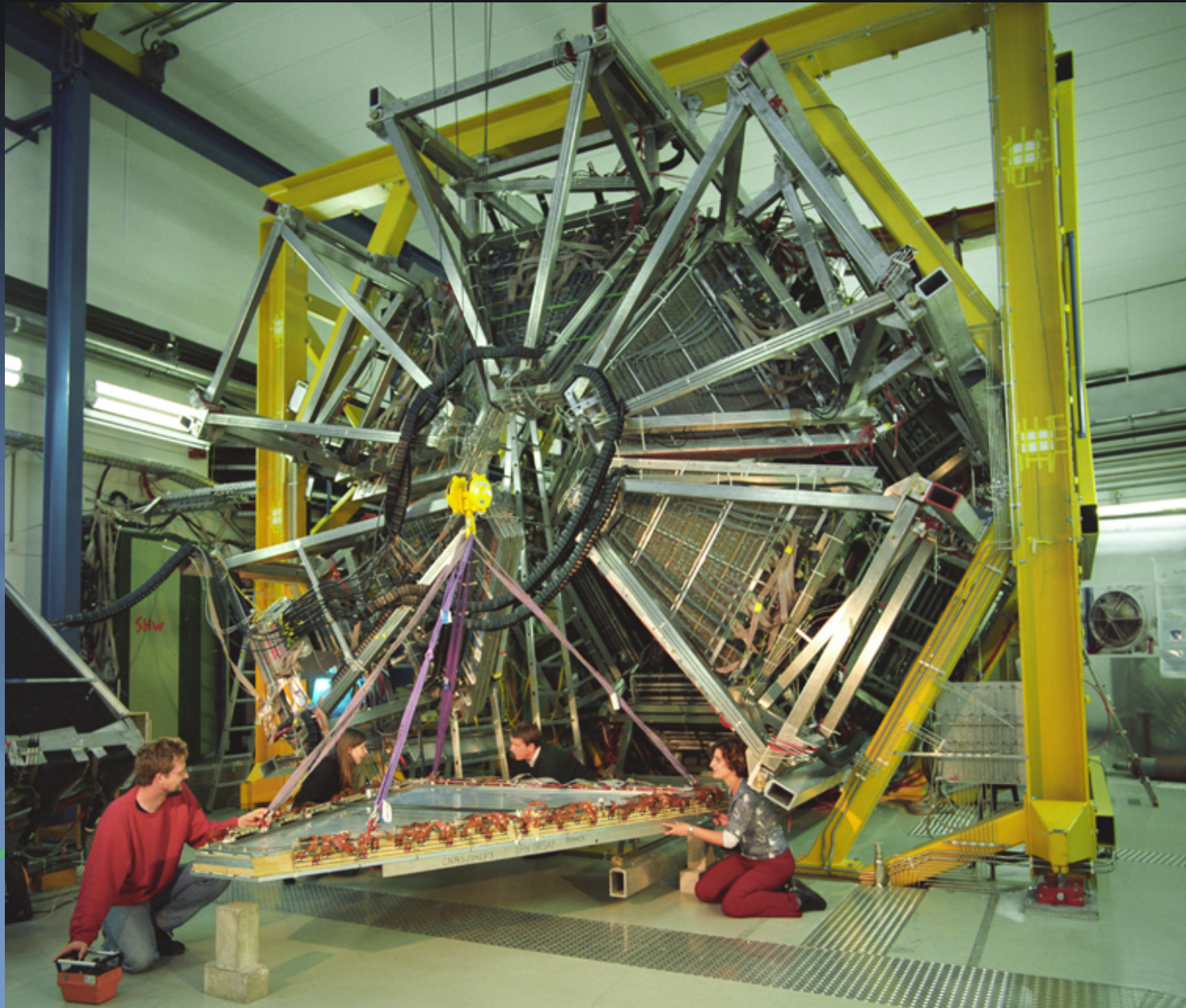
- **kinetic energy** determined via a calorimetric measurement

$$E_{kin} = (\gamma - 1)m_0c^2 \quad \gamma = \frac{1}{\sqrt{1 - \beta^2}}$$

- for  $Z=1$  the **mass** is extracted from  $E_{kin}$  and  $p$
- to determine  $Z$  (**particle charge**) a  $Z$ -sensitive variable is e.g. the ionization energy loss

$$\frac{dE}{dx} \propto \frac{z^2}{\beta^2} \ln(a \cdot \beta^2 \gamma^2) \quad a = \text{material-dependent constant}$$

# The HADES experiment @ GSI





# Rare ISotope INvestigation at GSI

